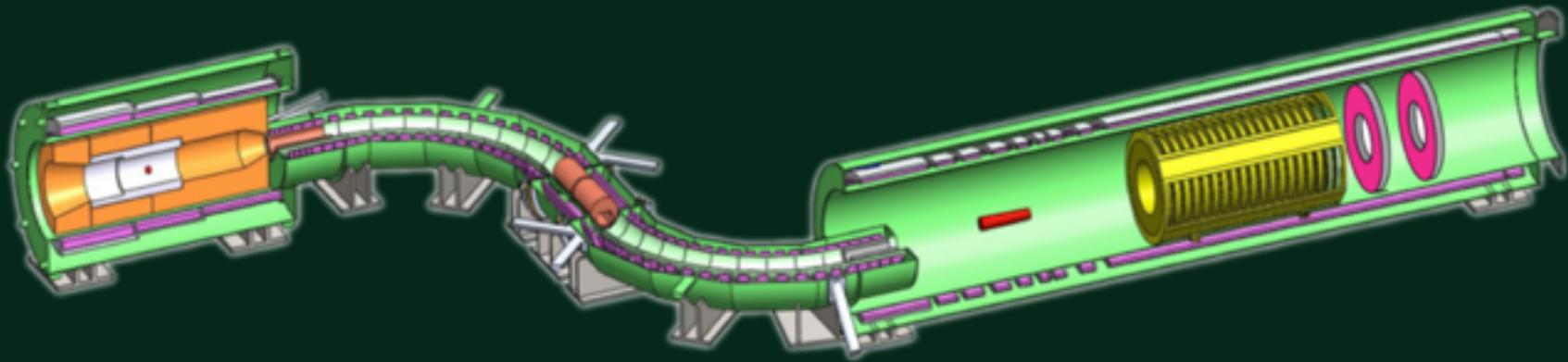


# Design, R&D and status of the crystal calorimeter for the Mu2e experiment



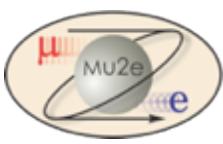
*Raffaella Donghia*

*Roma Tre University and LNF – INFN*

**May 12, 2016**  
**XVII LNF Spring School "Bruno Touschek"**

# Outline

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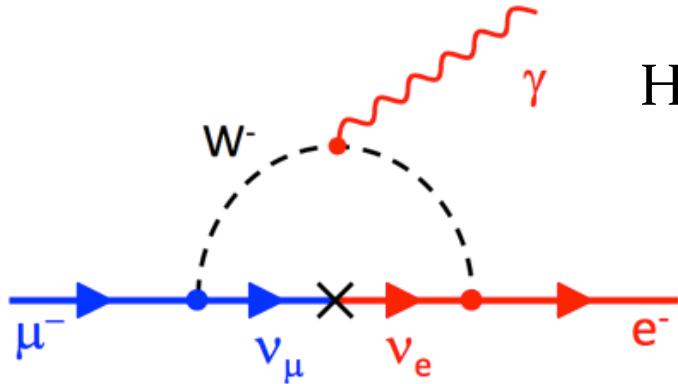
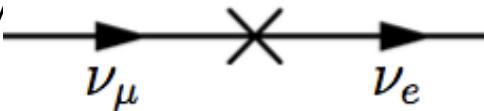


- CLFV
  - The muon sector
- The Mu<sub>2</sub>e apparatus
  - Detector Layout
- The crystal Calorimeter
  - Test results
- Conclusions

# Charged Lepton Flavor Violation

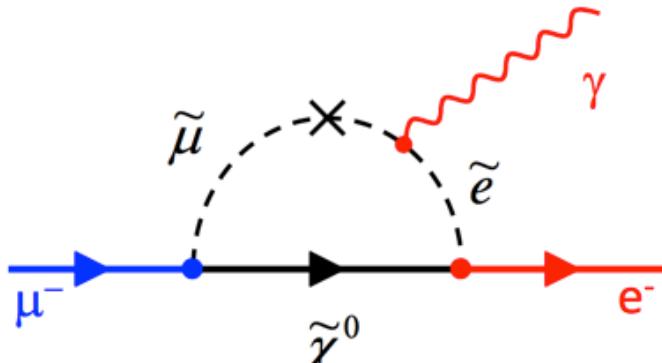
**Muon-to-electron conversion is a charged lepton flavor violating process (CLFV)**

Even in the SM, neutral LFV implies CLFV through neutrino mixing



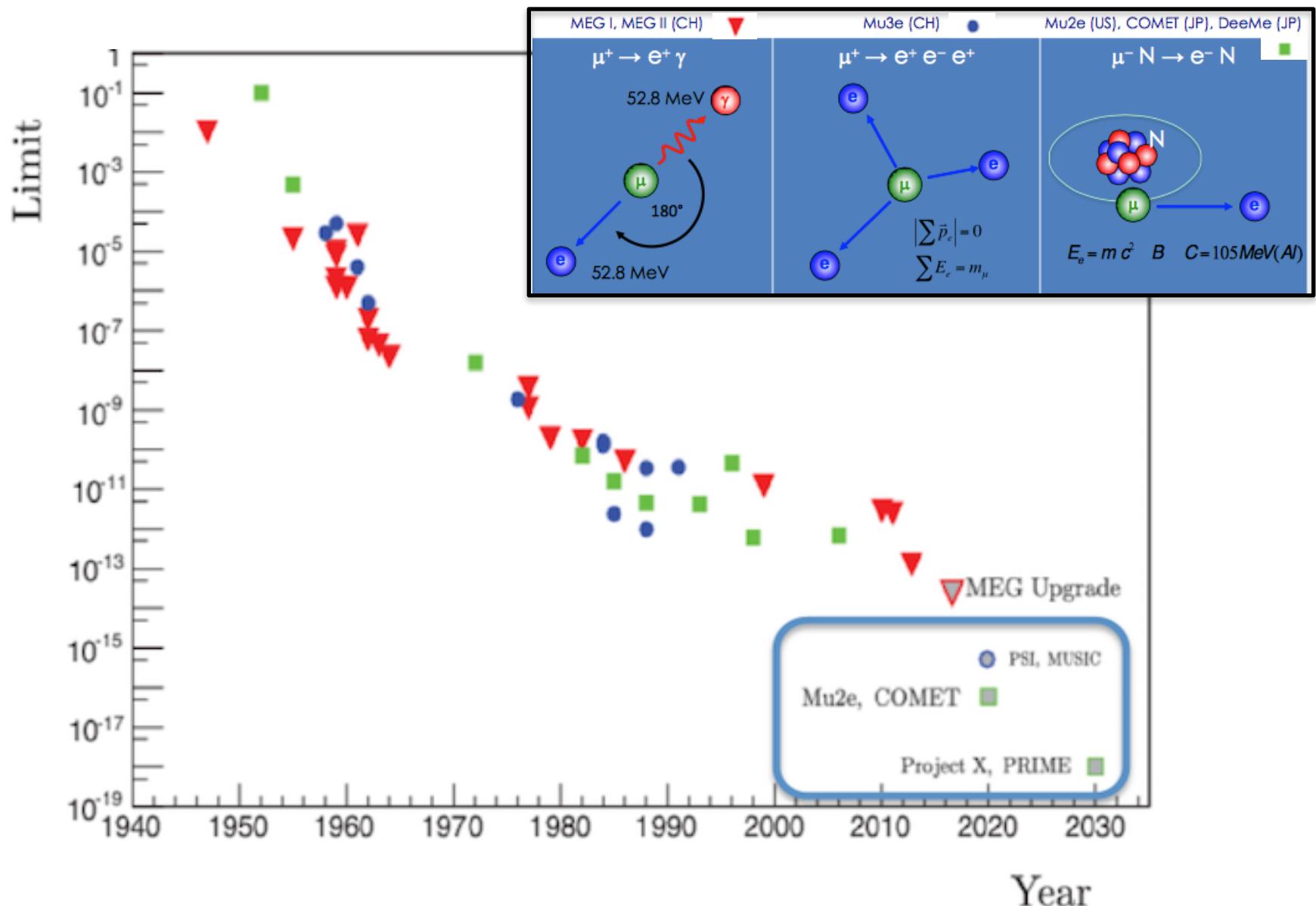
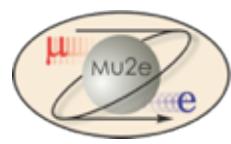
However, CLFV processes are strongly suppressed in the SM:  
 $\text{BR}(\mu \rightarrow e \gamma) < 10^{-54}$

New Physics can enhance CLFV rates to observable values



**Observation of CLFV: unambiguous sign of NP**

# Muon-CLFV history



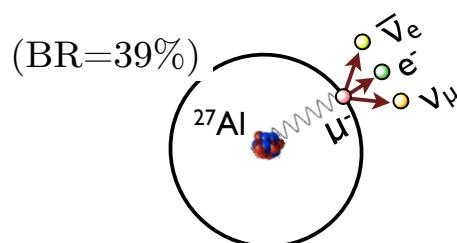
# Mu2e in one page



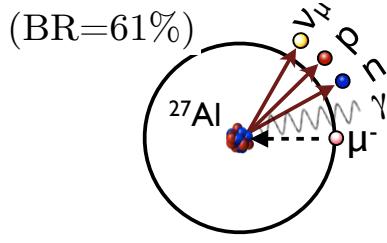
- Make muonic Al

- Low momentum  $\mu$  beam ( $< 100 \text{ MeV}/c$ )
- High intensity “pulsed” rate

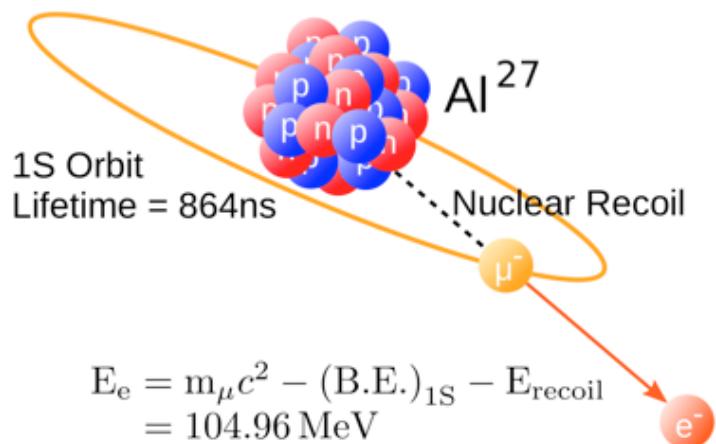
- Watch it decay:



a) Decay in Orbit (DIO)



b) Muon Capture



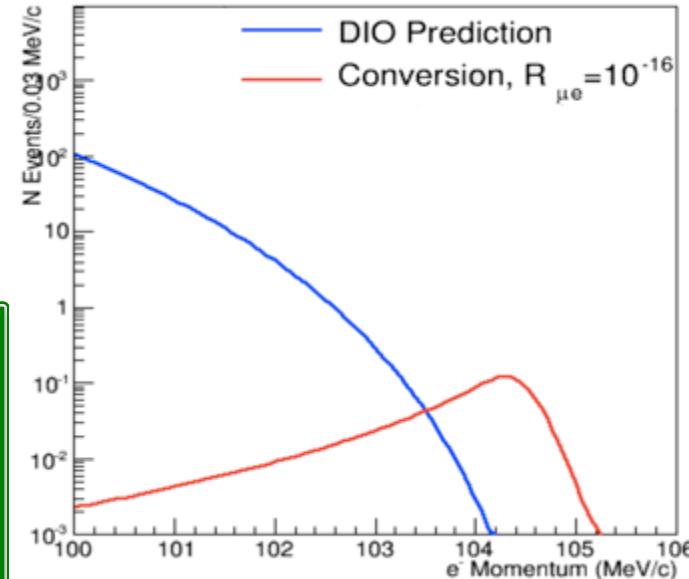
c) Conversion Process

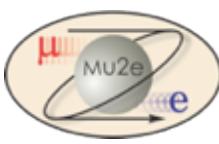
- Neutrinoless muon conversion

- Monoenergetic electrons  $E_e \sim 105 \text{ MeV}$
- At the endpoint of the DIO spectrum

Design sensitivity for a 3 year run (@ 90% C.L.):

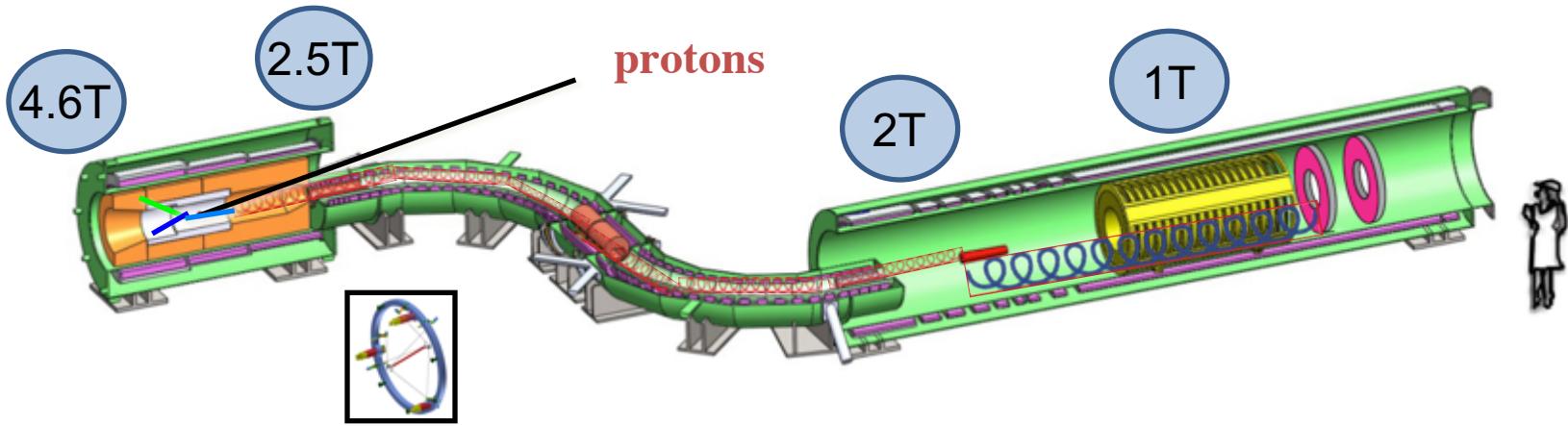
$$R_{\mu e} = \frac{\Gamma(\mu^- + A(Z, N) \rightarrow e^- + A(Z, N))}{\Gamma(\mu^- + A(Z, N) \rightarrow \nu_\mu + A(Z - 1, N))} < 6 \times 10^{-17}$$





## Production Target / Solenoid (PS)

- 8 GeV Proton beam strikes target, producing mostly pions
- Graded magnetic field contains backwards pions/muons and reflects slow forward pions/muons



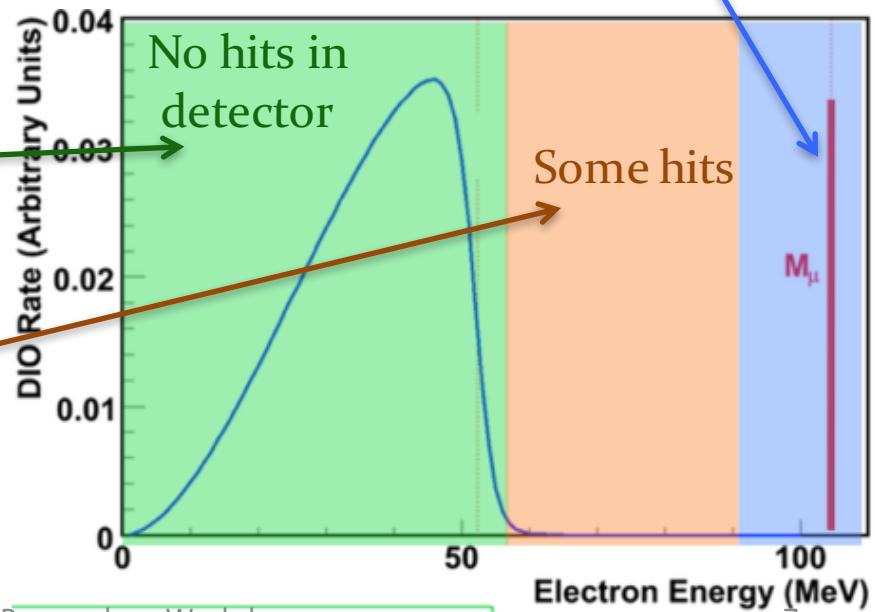
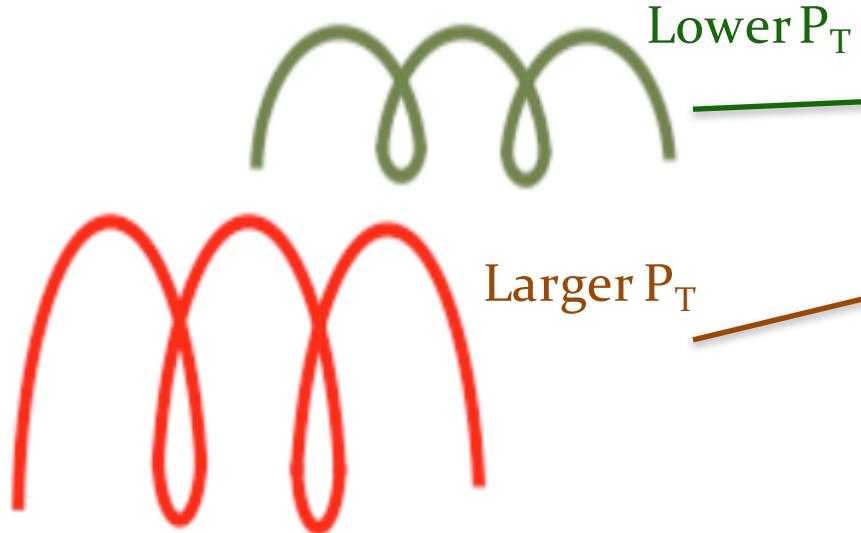
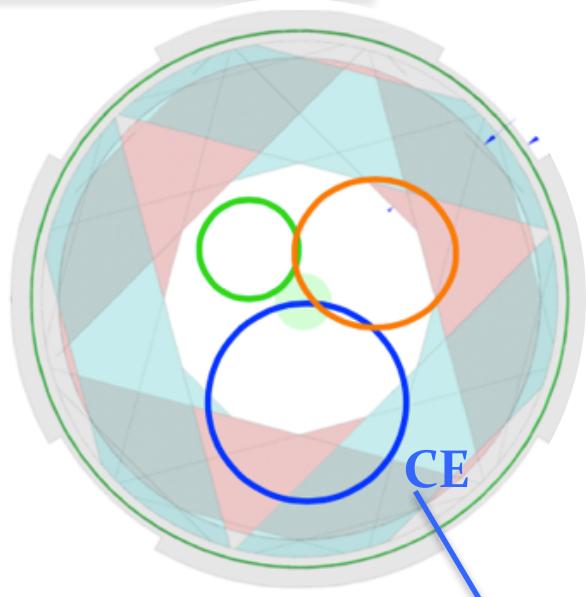
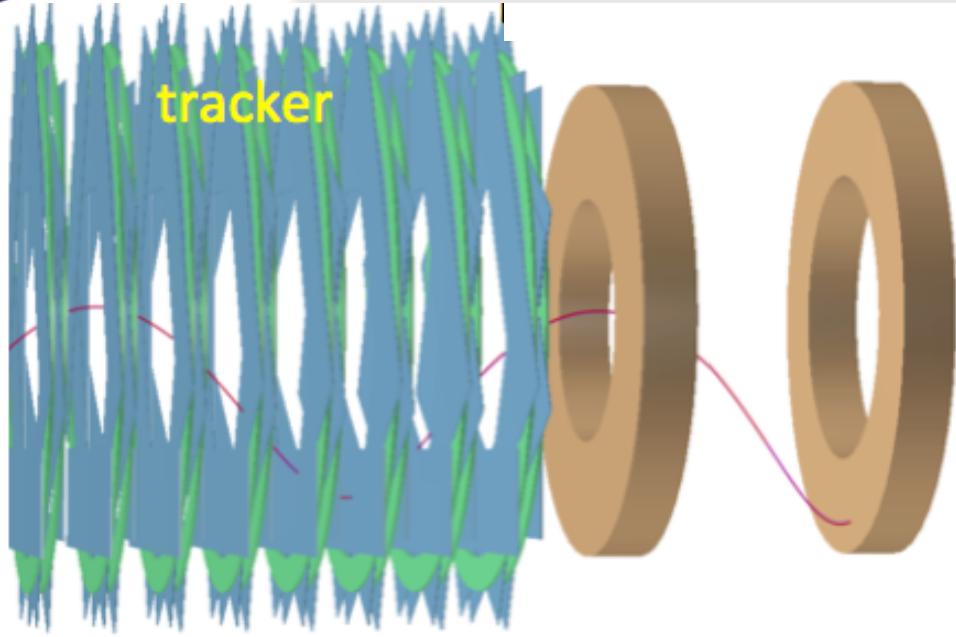
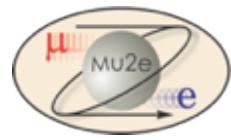
## Transport Solenoid (TS)

- Selects low momentum, negative muons
- Antiproton absorber in the mid-section

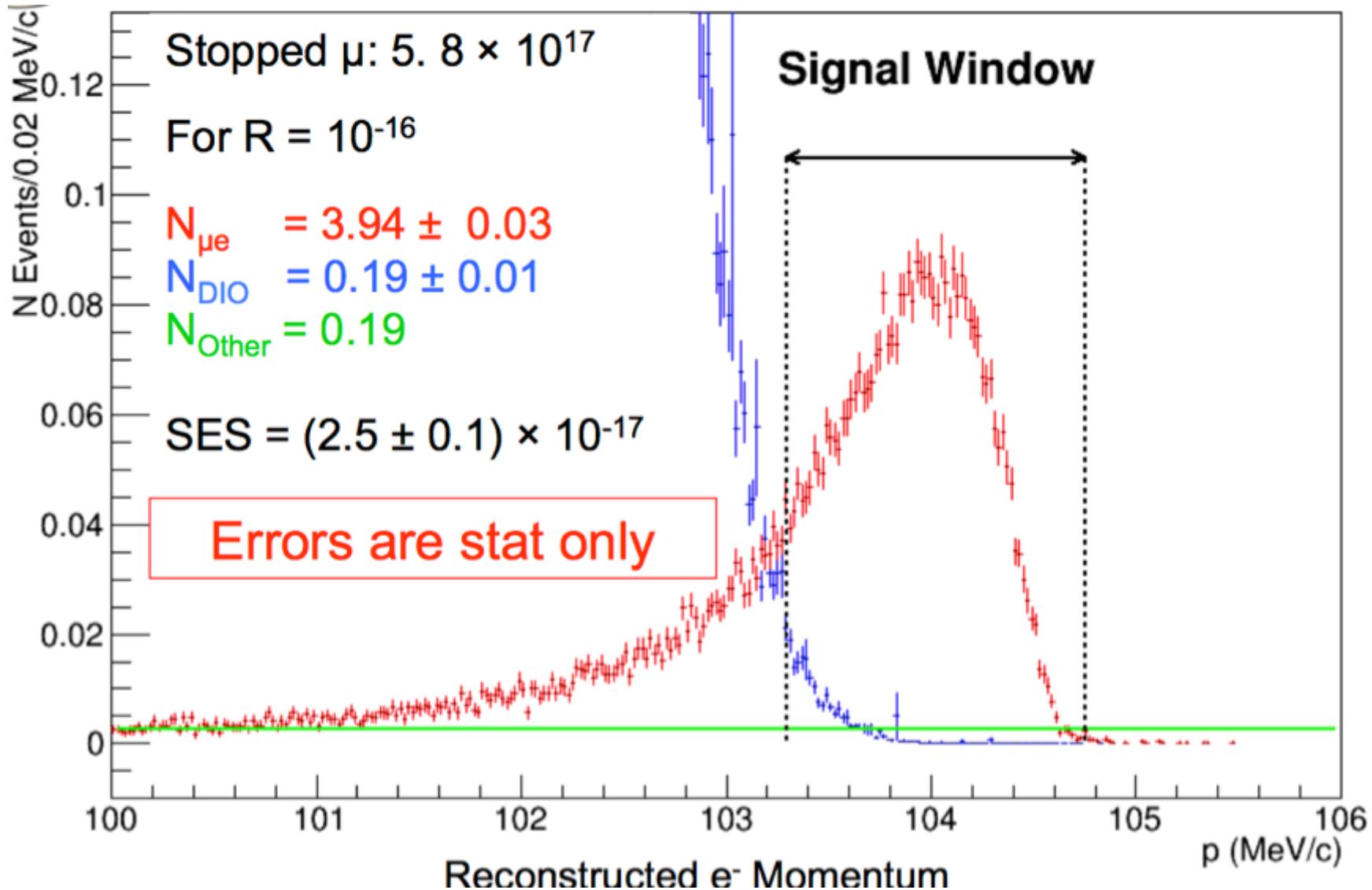
## Detector Solenoid (DS)

- Capture muons on Al target
- Measure momentum in tracker (with a resolution better than 120 keV @ 100 MeV) and energy in calorimeter
- CRV to veto Cosmic Rays event

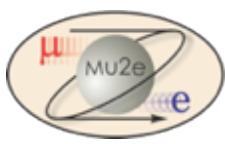
# How do you reach a SES of $2.5 \times 10^{-17}$ ?



# Simulated performances



# Calorimeter Requirements

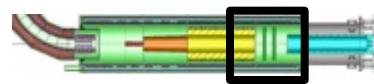


In order to add redundancy to this “super-rare” search, the calorimeter has to add complementarity qualities to the tracker:

- Large acceptance for CE
- Independent measurement of **energy / time / position**
- **PID capabilities**
- Independent trigger
- “seeds” to improve track finding efficiency at high occupancy

**In order to do so the calorimeter should provide:**

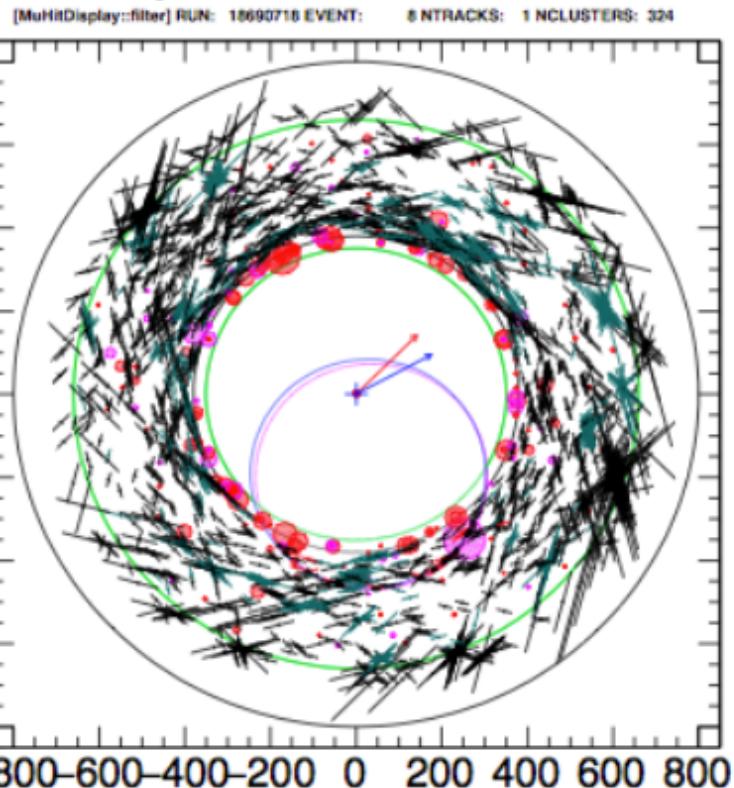
- **energy resolution  $\sigma_E/E$  of O(5 %)**
- **timing resolution  $\sigma_{(t)} < 500$  ps**
- position resolution < 1 cm
- Crystals survive a radiation dose of 100krad and a neutron fluence of  $10^{12} n/cm^2$
- Photo-sensors survive a neutron fluence of  $3 \times 10^{11} n_1 MeV/cm^2$



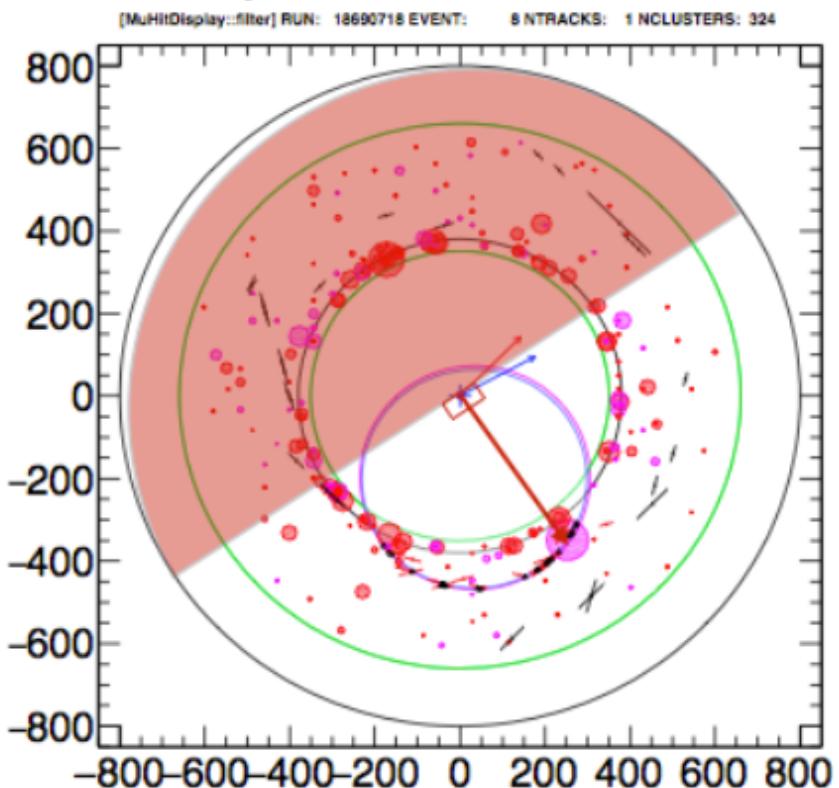
# Track seeding

Speed and efficiency of tracker reconstruction is improved by selecting the calorimeter clusters and the tracker hits comparable with the time ( $|\Delta t| < 50$  ns) and azimuthal angle of calorimeter clusters

**ce + spurious hits: no selection**

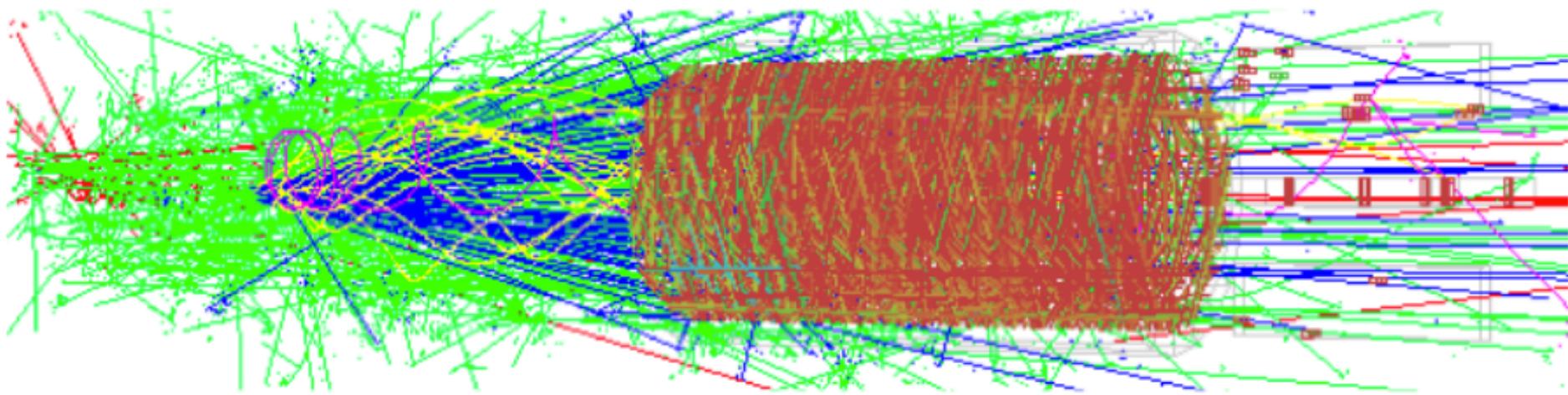


**ce + spurious hits: calo selection**

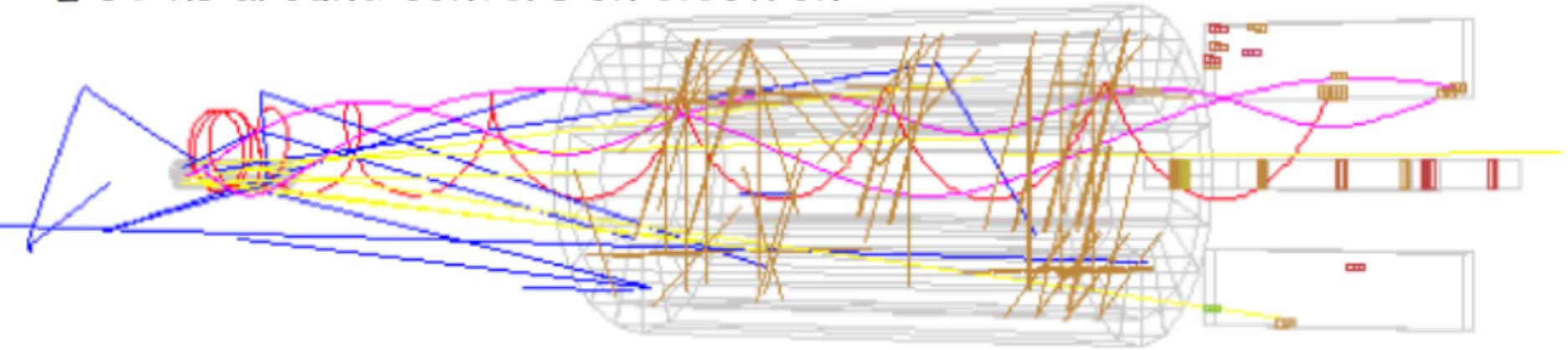


# A single micro bunch event

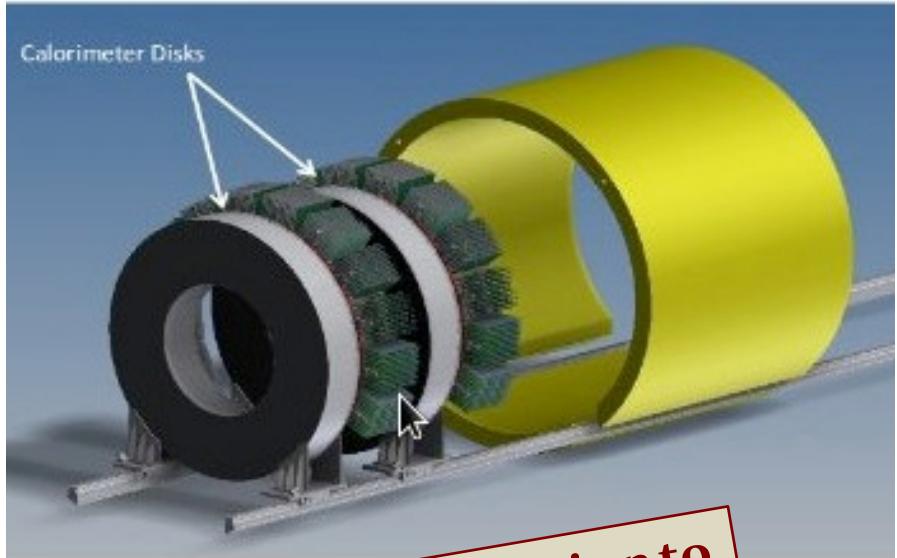
500 - 1695 ns window



$\pm 50$  ns around conversion electron

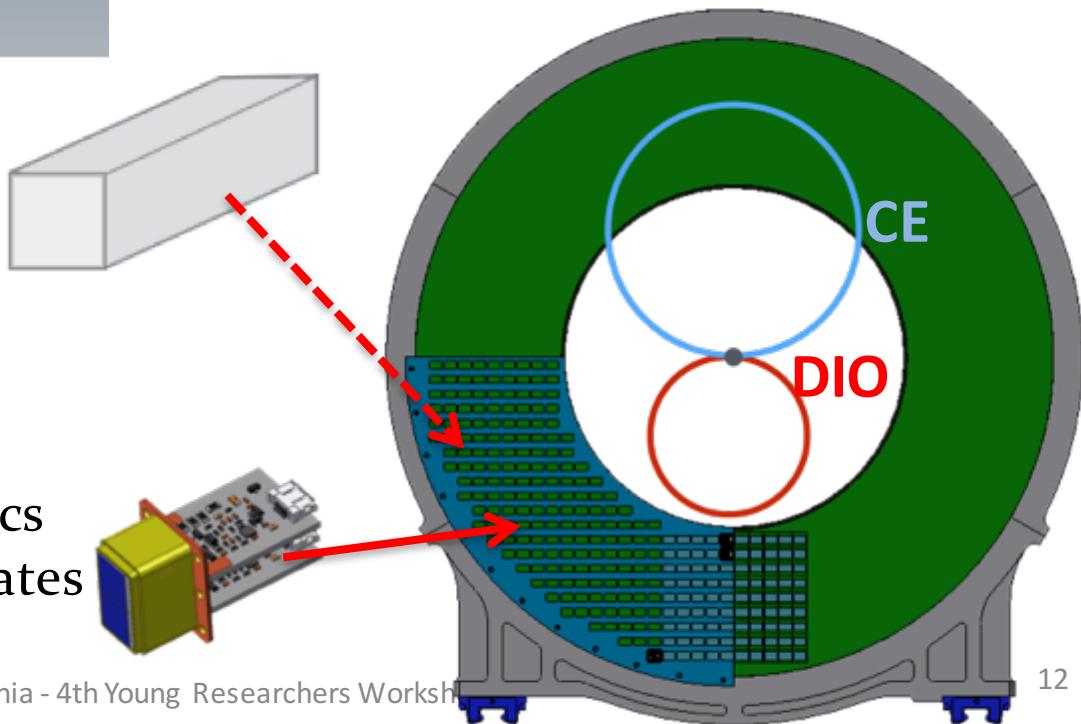


# Calorimeter



**Significant contribution to  
the calorimeter design  
from the INFN group**

- 2 annular disks with 1400 square crystals ( $34 \times 34 \times 200$ ) mm<sup>3</sup>
- $R_{IN} = 351$  mm,  $R_{OUT} = 660$  mm,  
Depth =  $10 X_0$  (200 mm)
- Distance 70 cm



- 2 SiPMs/crystal (2800 total)
- Analog FEE and digital electronics located in near-by electronics crates

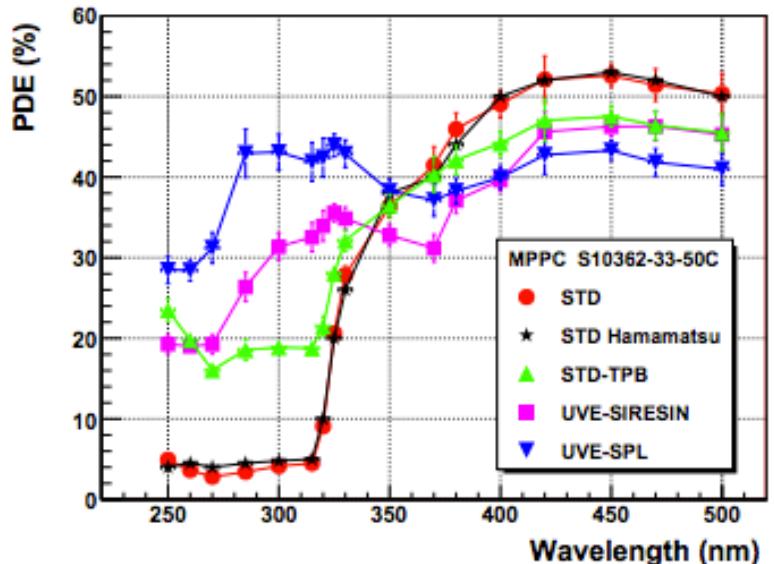
# Crystals and Sensors choice

	LYSO	BaF <sub>2</sub>	CsI
Radiation Length X <sub>0</sub> [cm]	1.14	2.03	<b>1.86</b>
Light Yield [% NaI(Tl)]	75	4/36	<b>3.6</b>
Decay Time[ns]	40	0.9/650	<b>30</b>
Photosensor	APD	RMD APD	<b>SiPM</b>
Wavelength [nm]	402	220/300	<b>310</b>

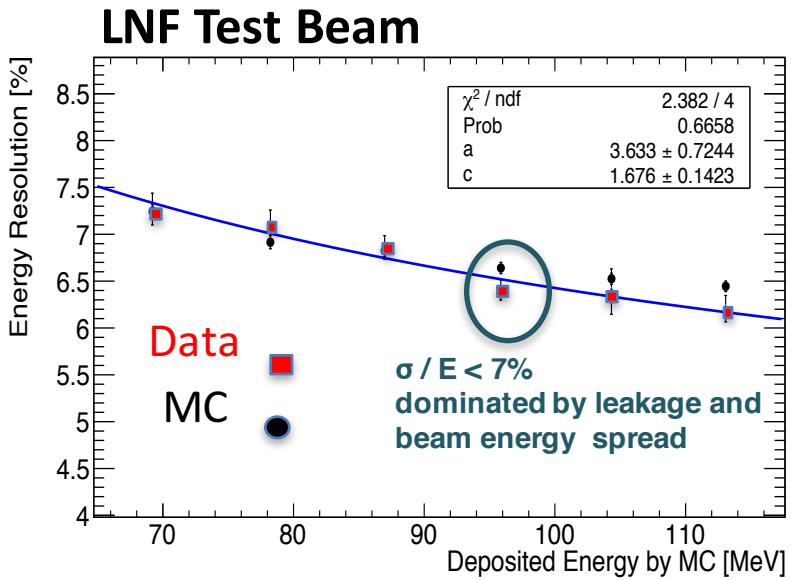


## CsI(pure)

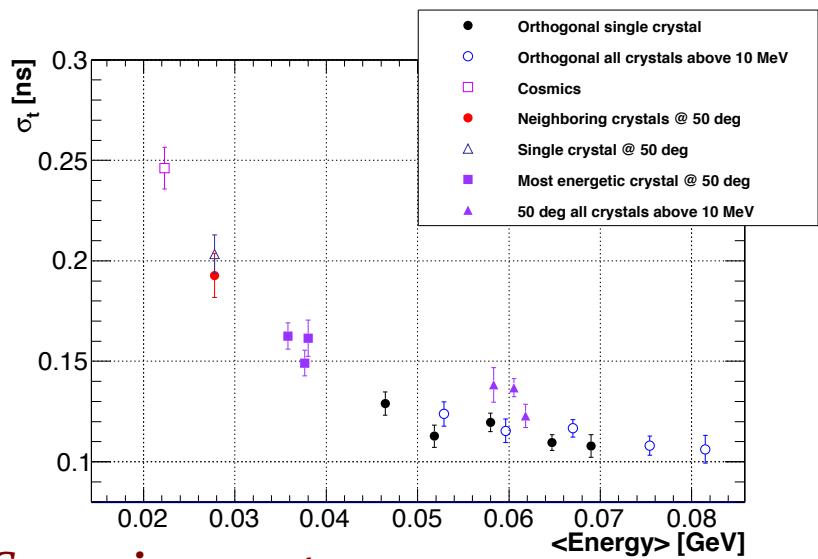
- Adequate radiation hardness
- Slightly hygroscopic
- 30 ns emission time, small slow component.
- Emits @ 310 nm.
- Comparable LY of fast component of BaF<sub>2</sub>.
- Lower cost (6-8 \$/cc)



# Meeting the requirements (1)



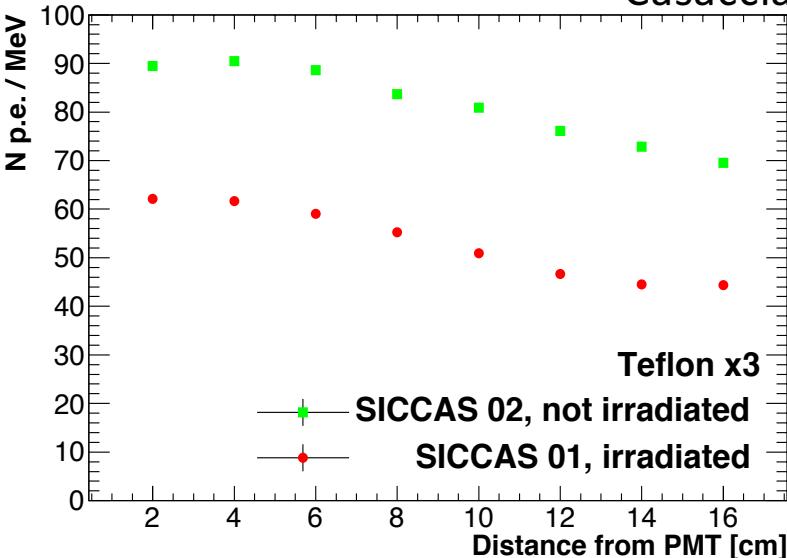
- Simulation performed as a function of LY and many other variables → **CsI+SIPM match requirements**
- Test beam with  $e^-$  @ BTF, LNF 80 to 130 MeV 3x3 array of  $30 \times 30 \times 200 \text{ mm}^3$  CsI + MPPC used
- Good energy ( 7%) and timing ( 110 ps) resolution



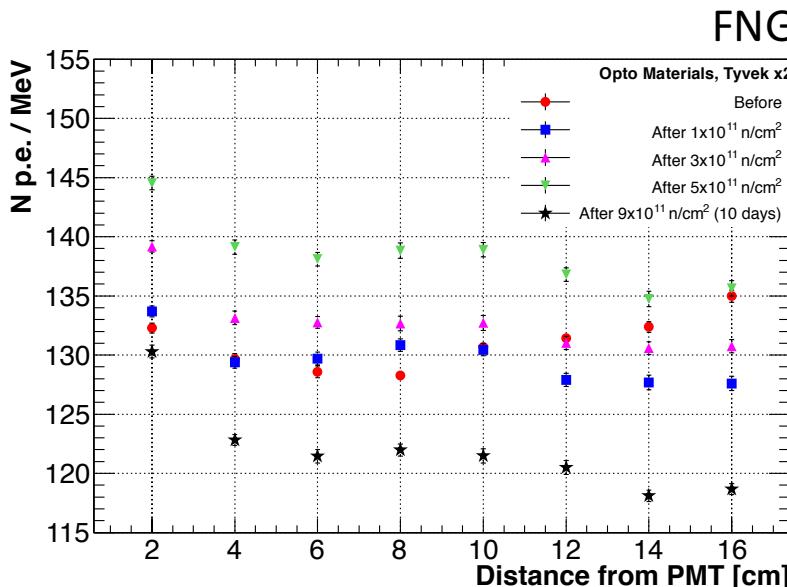
**Exp tests → Matching EMC requirements**

# Meeting the requirements (2)

- **Crystals have been tested up to 100 krad and  $10^{12}$  n/cm<sup>2</sup> with 14 MeV neutrons.**
  - No major issues/damages observed (LY drop 40% @ 100 krad)
  - Radiation hardness will be part of our QA test procedure
  - Effect of thermal neutrons on Radiation Induced Current being investigated

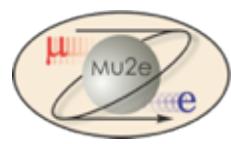


- **Photosensors have been tested up to 20 krad and  $3 \times 10^{11}$  n/cm<sup>2</sup>.**
  - No problems with dose
  - With neutrons, sensors are still working but leakage current increases to high values
  - Cooling photosensors to 0 C required.



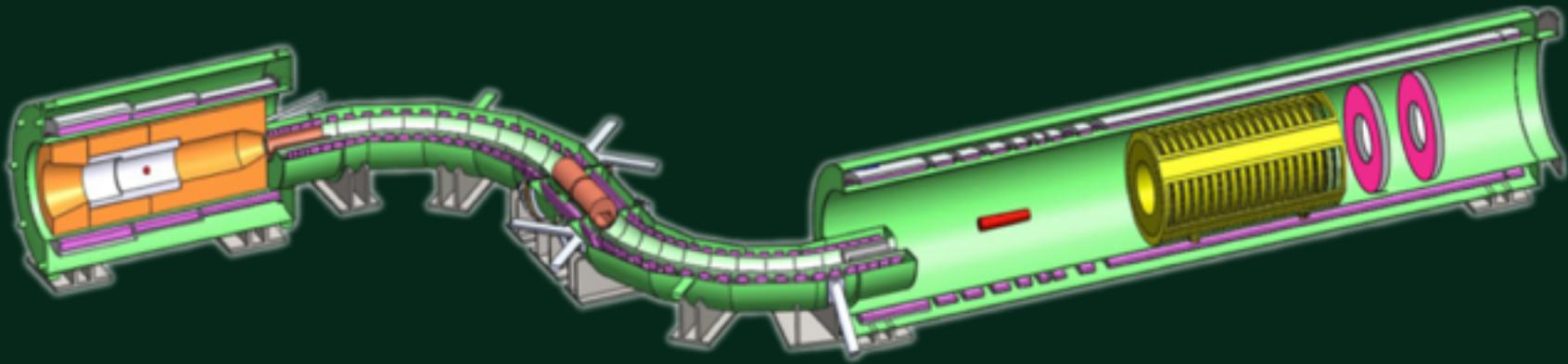
# Conclusions

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- Mu2e is a CLVF first-class experiment looking for NP BSM with high complementarity to other programs while increasing reach and diversification in models testing
- Mu2e will improve previous conversion experiment of 4 orders of magnitude and probe mass scales up to hundreds of TeV.
- < 10 years Timeline for completion of first phase.
- Mu2e has completed the CD-2 and CD3 for the long lead items
  - Construction of the solenoids will start next year.
  - Detector Review in spring to freeze detector with CD3 in summer 2016
  - Construction period 2017-2019, followed by installation in 2019 - 2020
- Mu2e-2 phase planned, increasing (x 10) intensity and sensitivity!

# Thank You!



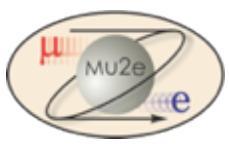
*Raffaella Donghia*

*Roma Tre University and LNF - INFN*

XVII LNF Spring School "Bruno Touschek"

# backup

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# Why muon conversion is unique?

**Muon conversion is a unique probe for BSM:**

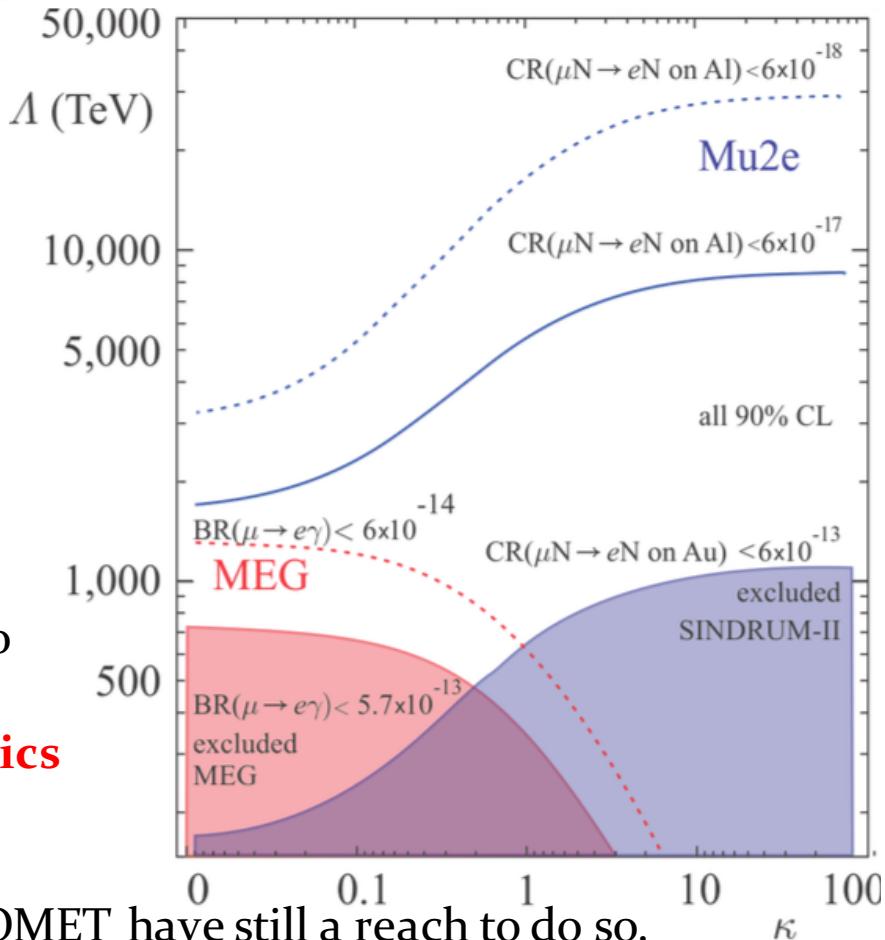
◆ **Broad discovery sensitivity across all models:**

- Sensitivity to the same physics of MEG but with better mass reach
- Sensitivity to physics that MEG is not
- If MEG observes a signal, MU2E/COMET do it with improved statistics.

**Ratio of the BR allows to pin-down physics model**

- If MEG does not observe a signal, MU2E/COMET have still a reach to do so.

◆ **Sensitivity to  $\lambda$  (mass scale) up to hundreds of TeV beyond any current existing accelerator**



# Possible upgrade

Signal ?

Yes

Let's party!

Change the target to study the  
underlying NP

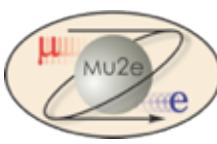
Redesign muon transport and  
detector

No

Higher rates, background  
must decrease to measure R  
at  $10^{-18}$

Redesign muon transport  
and detector

# CLFV in the muon sector



MEG I, MEG II (CH)

Mu3e (CH)

Mu2e (US), COMET (JP), DeeMe (JP)

$$\mu^+ \rightarrow e^+ \gamma$$

52.8 MeV



$\gamma$

180°

52.8 MeV

180°

52.8 MeV

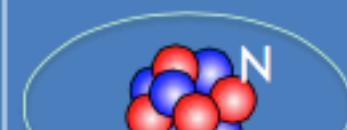
$$\mu^+ \rightarrow e^+ e^- e^+$$



$$|\sum \vec{p}_e| = 0$$

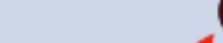
$$\sum E_e = m_\mu$$

$$\mu^- N \rightarrow e^- N$$



$$E_e = m c^2 \quad B \quad C = 105 \text{ MeV}(A)$$

< 52.8 MeV



$\gamma$

52.8 MeV

52.8 MeV

52.8 MeV

52.8 MeV

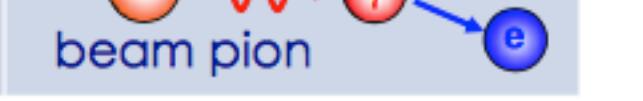
< 52.8 MeV

52.8 MeV

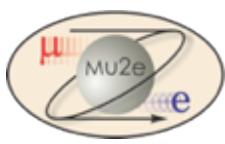
52.8 MeV

52.8 MeV

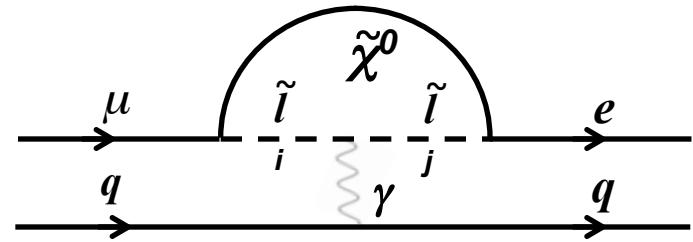
$$E_e < m$$



# Muon conversion



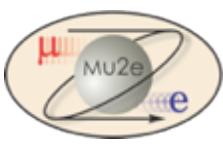
- The Mu2e experiment searches for muon-to-electron conversion in the coulomb field of a nucleus:  $\mu^- Al \rightarrow e^- Al$
- $\mu$ -e is a CLFV process, similar but complementary to other CLFV processes as  $\mu \rightarrow e\gamma$  and  $\mu \rightarrow 3e$
- It is a process strongly suppressed in the SM, but NP could enhance CLFV rates to observable values
  - **SO(10) SUSY** (L. Calibbi et al., hep-ph/0605139)



- **Higgs triplet model** (M. Kakizaki et al., PLB 566)
- **Littlest Higgs with T-parity** (M. Blanke et al., ActaPhys.Polon.B 41:657)
- **Scalar Leptoquarks** (J.M. Arnold et al., Phys.Rev.D 88 035009)
- **Left-Right symmetric model** (C.H. Lee et al., Phys.Rev.D 88, 093010)

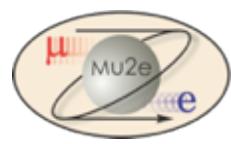
# List of Backgrounds to fight

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- **Muon decay in orbit (DIO)**
- **Radiative pion capture (RPC)**  
 $\pi^- N \rightarrow \gamma N'$ ,  $\gamma \rightarrow e^+ e^-$  and  $\pi^- N \rightarrow e^+ e^- N'$
- Antiprotons: produce pions when they annihilate in the target .. antiprotons are negative and they can be slow!
- Pion/muon decay in flight
- Electrons from beam
- **Cosmic rays**
- ...

# DIO background

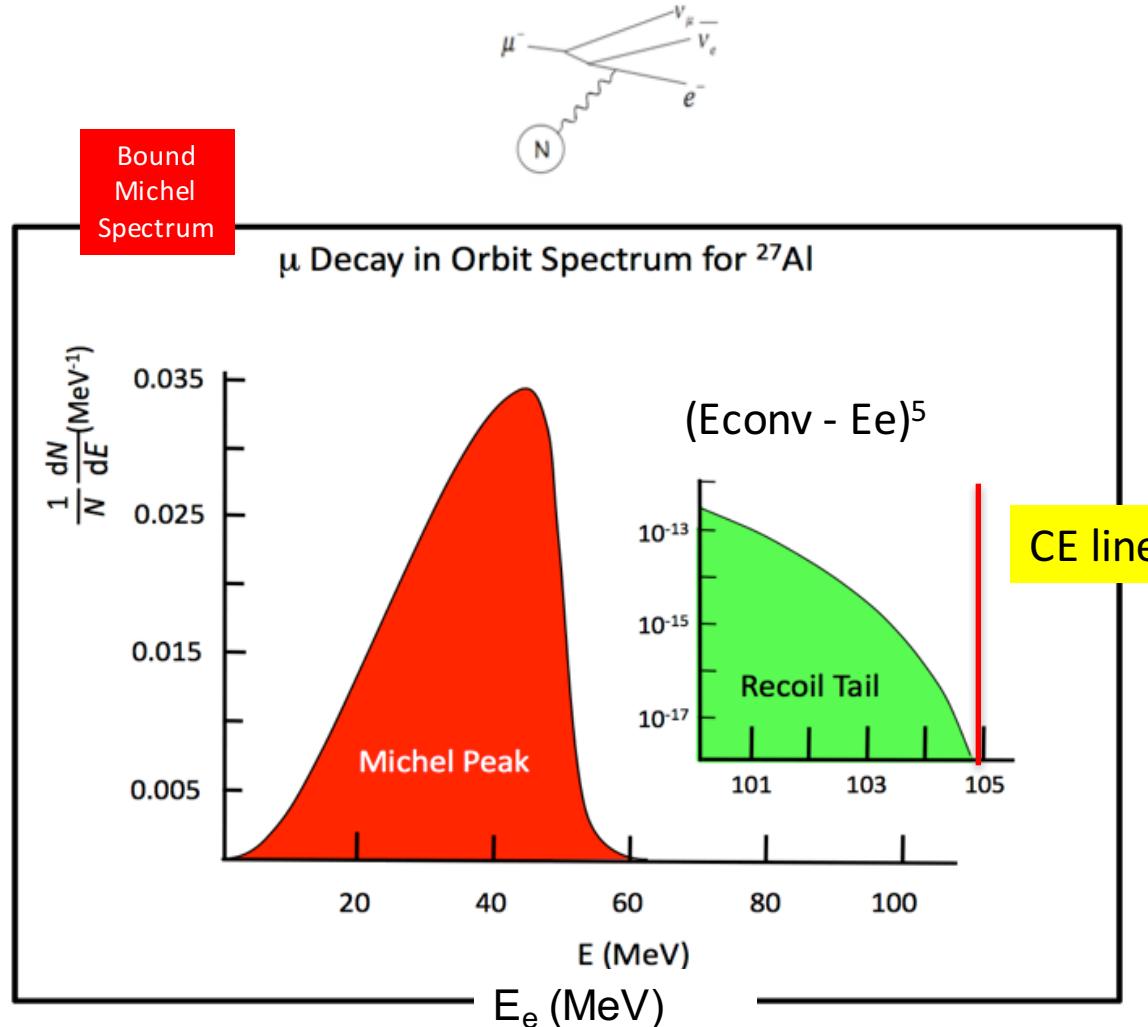


- The DIO background is the most difficult one.

□ Electron energy distribution from the decay of bound muons is a (modified) Michel spectrum:

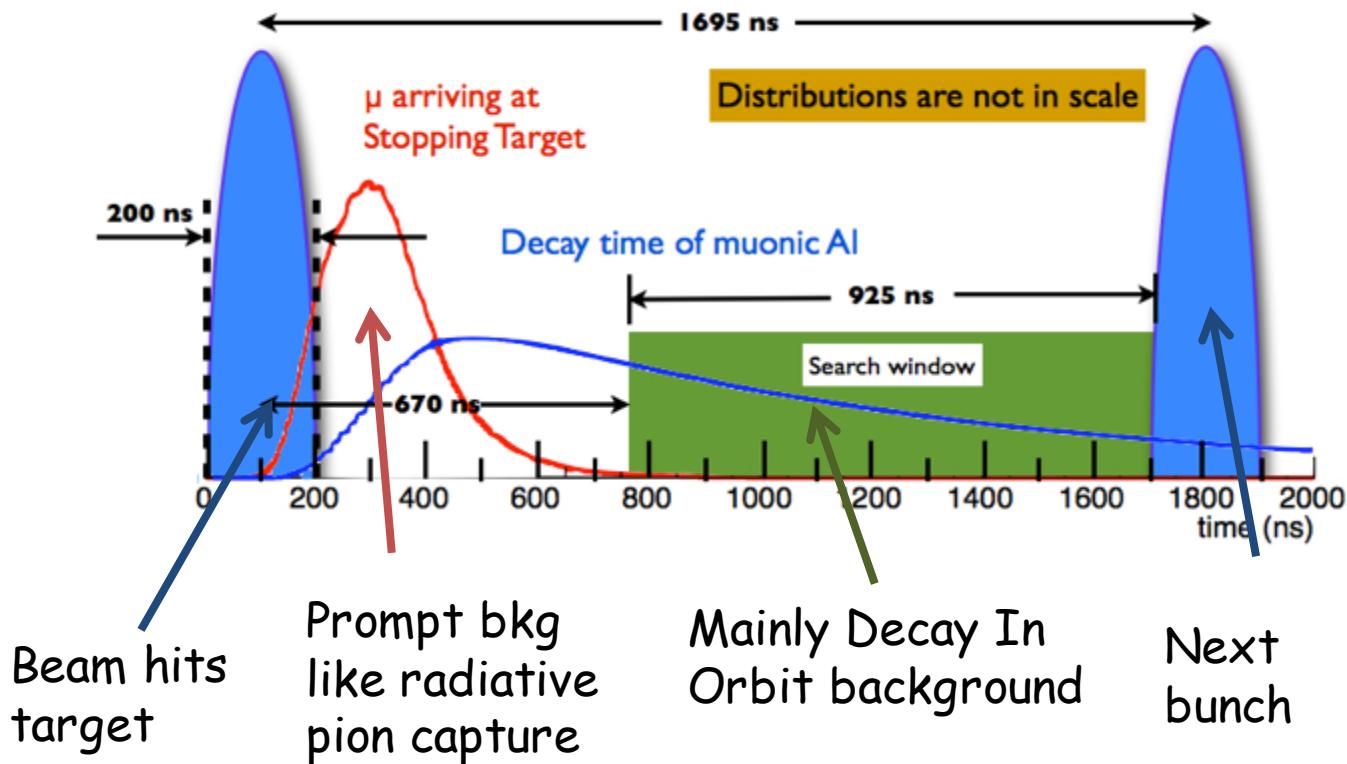
→ Presence of atomic nucleus and momentum transfer create a recoil tail with a fast falling slope close to the endpoint

→ To separate DIO endpoint from CE line we need a high Resolution Spectrometer



Czarnecki et al., Phys. Rev. D 84, 013006 (2011)  
arXiv:1106.4756v2

# Beam structure → prompt background



## ❑ Use the fact that muonic atomic lifetime >> prompt background

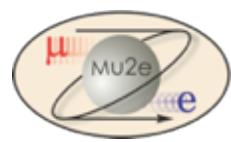
Need a pulsed beam to wait for prompt background to reach acceptable levels  
 → Fermilab provides the beam we need !

## ❑ OUT of time protons are also a problem.

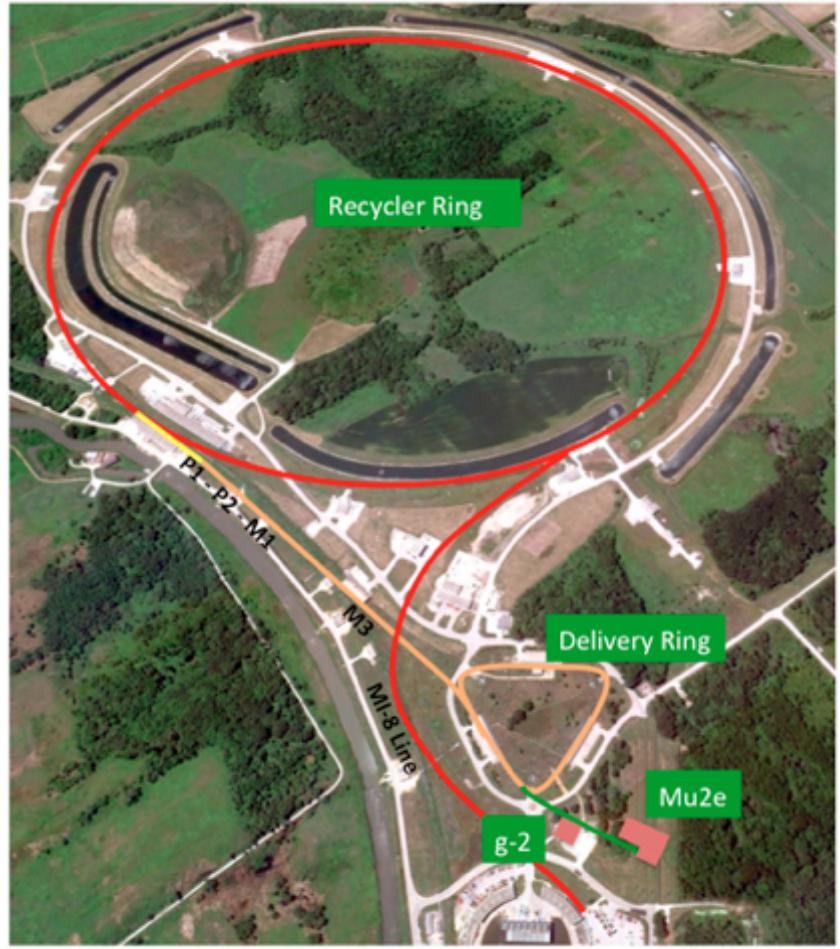
To keep associated background low we need proton extinction of  $10^{-10}$  :

proton extinction (between pulses) → # protons out of beam/# protons in pulse

# Accelerator Scheme & Proton extinction



- Booster: batch of  $4 \times 10^{12}$  protons every 1/15<sup>th</sup> second
- Booster “batch” is injected into the Recycler ring and re-bunched into 4 bunches
- These are extracted one at a time to the Delivery ring
- As a bunch circulates, protons are extracted to produce the desired beam structure → **bunches of  $\sim 3 \times 10^7$  protons each, separated by 1.7  $\mu$ s**



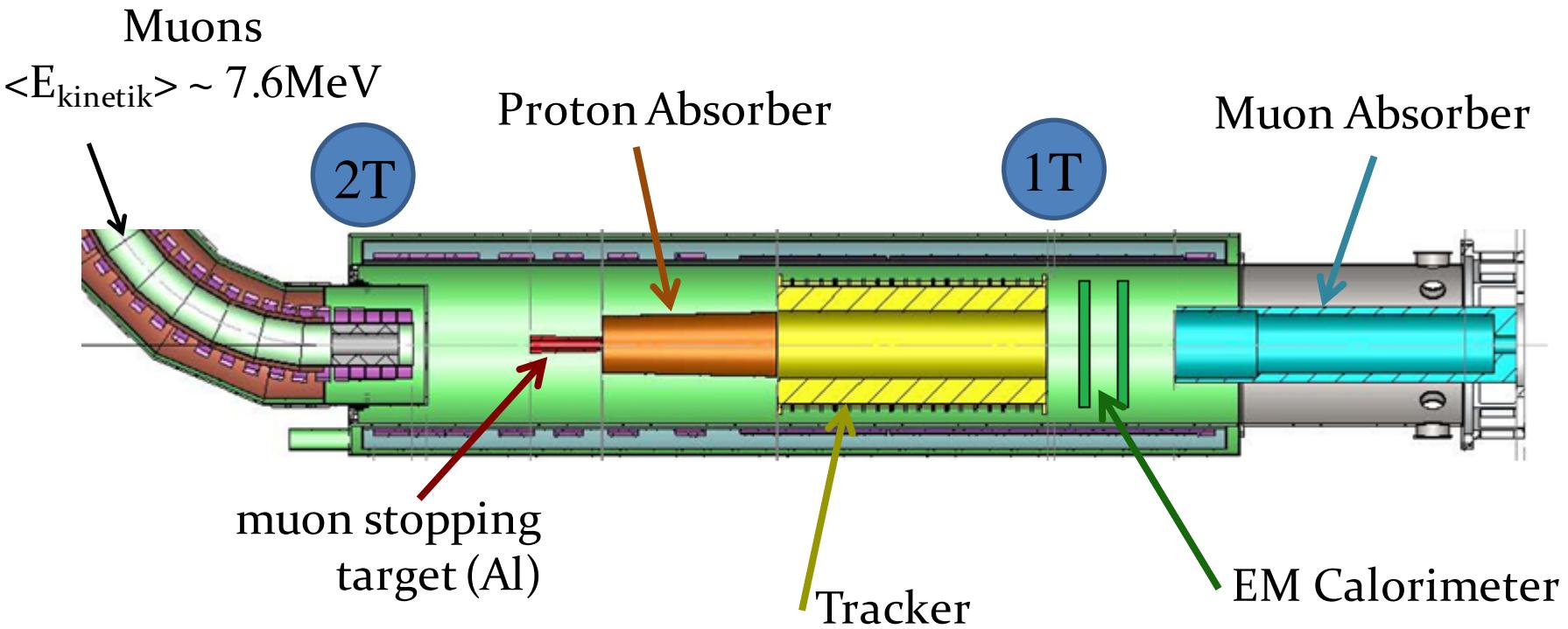
## Proton Extinction

*achieving  $10^{-10}$  is hard; normally get  $10^{-2} - 10^{-3}$*

- Internal (momentum scraping) and bunch formation in Accumulator
- External: oscillating (AC) dipole

**Calculations based on accelerator models  
That take into account collective effects  
Shows that this combination gets  $\sim 10^{-12}$**

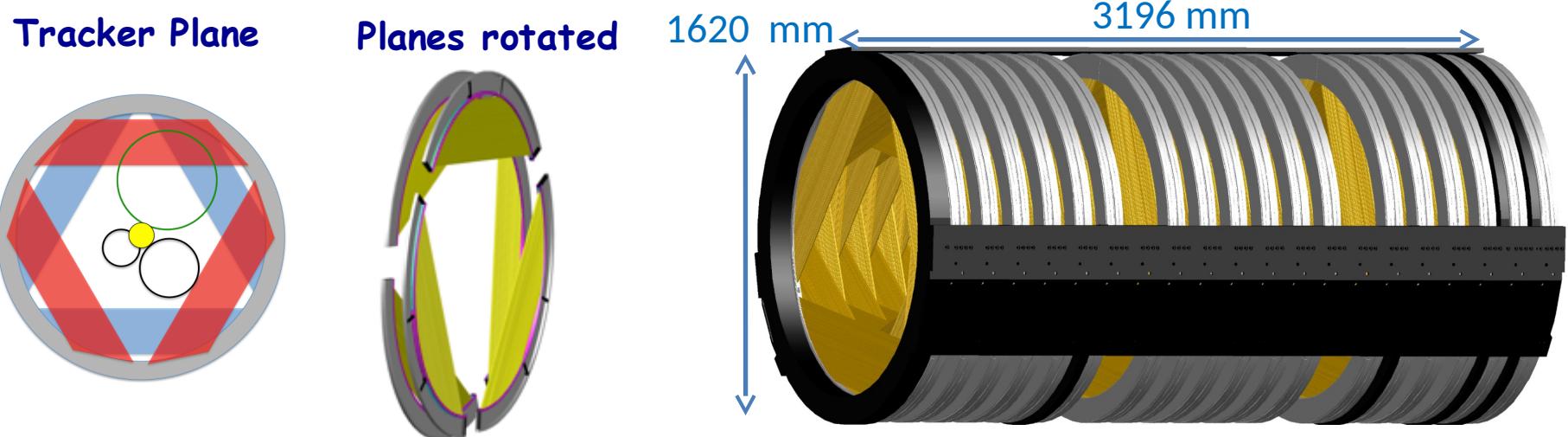
# The Detector Solenoid



- Graded field “reflects” downstream a fraction of conversion electrons emitted upstream (isotropic process)
- For the sensitivity goal →  $\sim 6 \times 10^{17}$  stopped muons for 3 years run  
→  $10^{10}$  stopped muon/s (10 GHz)

# Tracker system

- Tracker is a low mass straw drift tubes design with tubes transverse to secondary beam
- 15  $\mu\text{m}$  thick straw walls, 5 mm diameter, dual-ended readout, length 430 – 1120 mm.
- It must operate in vacuum
- $\sim 20000$  tubes arranged in planes on stations,
- The tracker has 18 stations.
- Tracking at high radius ensures operability: beam flash produces a lot of low momentum particles, large DIO background. Most of background miss the tracker.



# Calorimeter System (1)

## Calorimeter requirements:

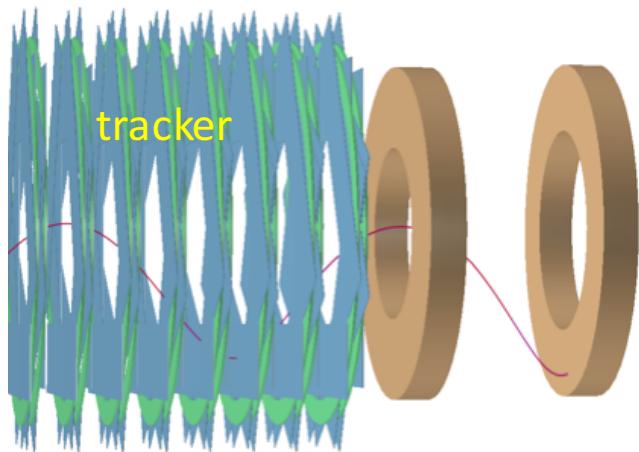
- Particle Identification to distinguish e/mu
- Seed for track pattern recognition
- Tracking independent trigger
- Work in 1 T field and  $10^{-4}$  Torr vacuum
- RadHard up to 30 krad,  $10^{12}$  n/cm<sup>2</sup>/year

## Calorimeter choice:

High granularity crystal based calorimeter with:

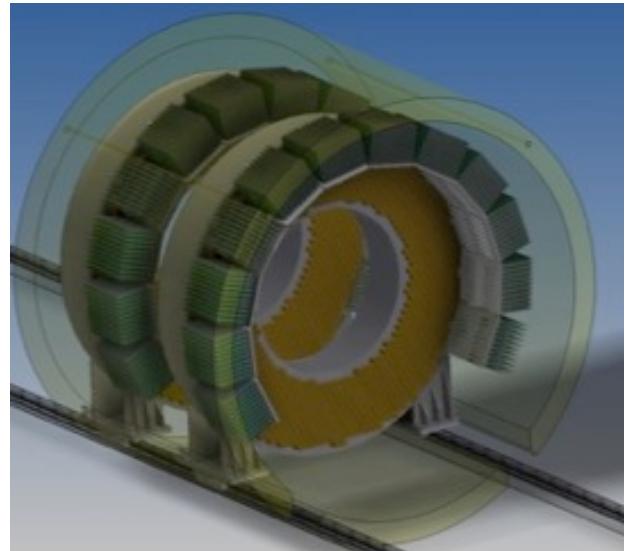
- $\sigma/E$  of O(5%) and Time resolution < 500 ps
  - Position resolution of O(1 cm)
  - almost full acceptance
- for CE signal @ 100 MeV

Two disks separated  
by  $\frac{1}{2}$  wavelength (70 cm)



## Disk geometry

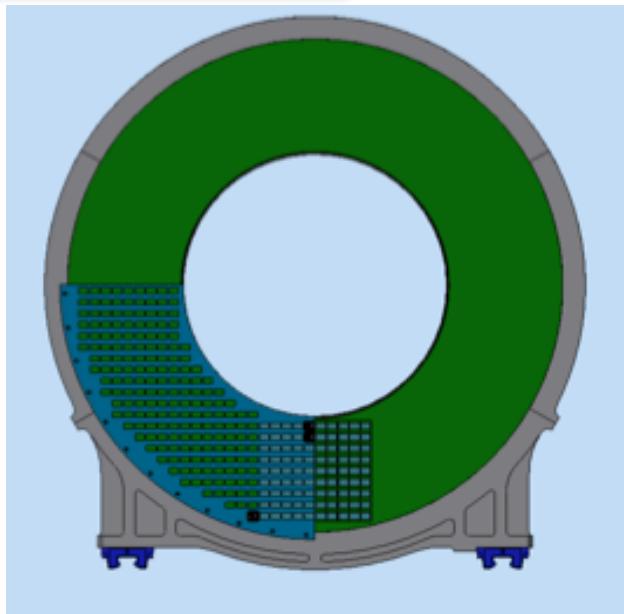
- Square crystals
- Charge symmetric, can measure  $\mu^- N \rightarrow e^+ N$



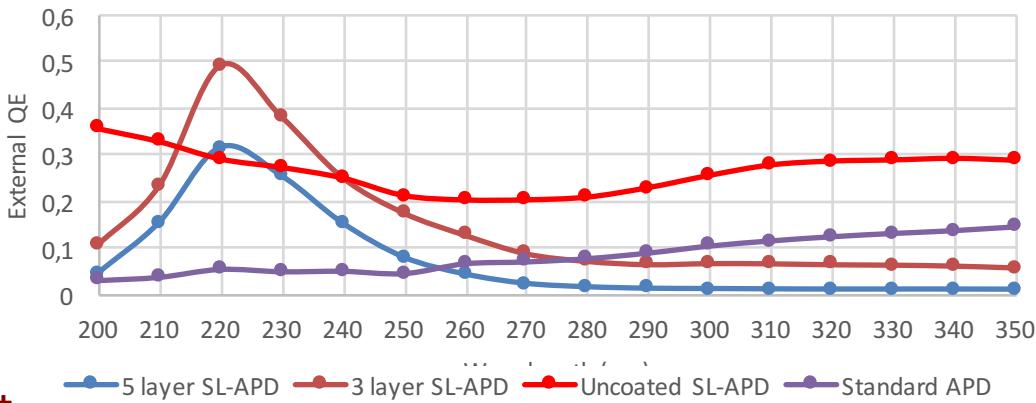
# Calorimeter System (2)

The Calorimeter consists of two disks with  
**1650 BaF<sub>2</sub> square crystals (30x30x200) mm<sup>3</sup>**

- R<sub>IN</sub> = 351 mm, R<sub>OUT</sub> = 660 mm, Depth = 10 X<sub>0</sub> (200 mm)
- Each crystal readout by two SL APDs (9x9 mm<sup>2</sup>)
- Analog FEE and digital electronics located on calo
- Radioactive source and laser systems provide absolute calibration and monitoring capability.



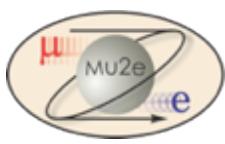
To reduce the slow BaF<sub>2</sub> component at higher wavelengths , a Caltech/JPL/RMD consortium formed to develop a RMD APD **into a super-lattice APD with high Q.E. @ 220 nm** that incorporates also **an Atomic Layer Deposition antireflection filter** to reduce efficiency for  $\lambda > 300$  nm.



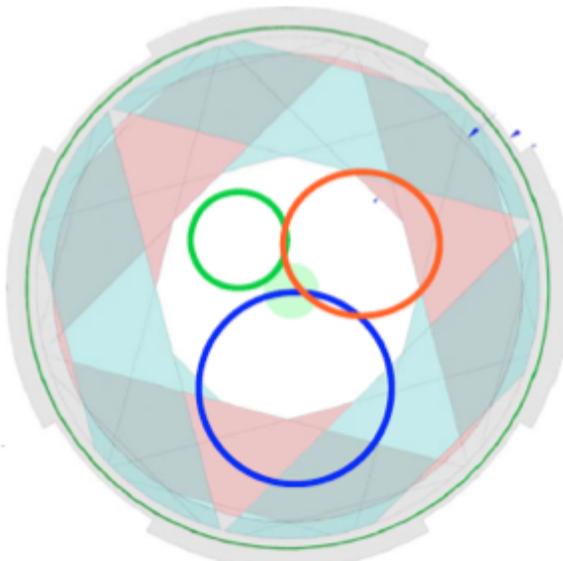
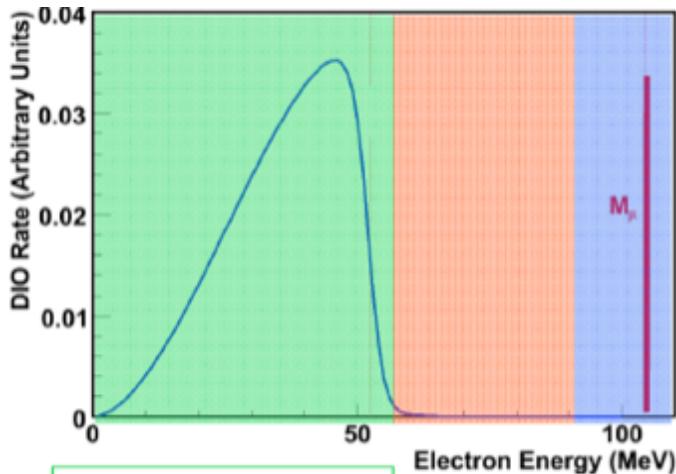
Prototypes with LYSO+APD, CsI+MPPC built  
Next one with BaF<sub>2</sub> + SL APDs in progress

Good progresses on FEE and mechanics

# Basic reconstruction scheme



## Reconstructable tracks



12/05/2016

Raffaell

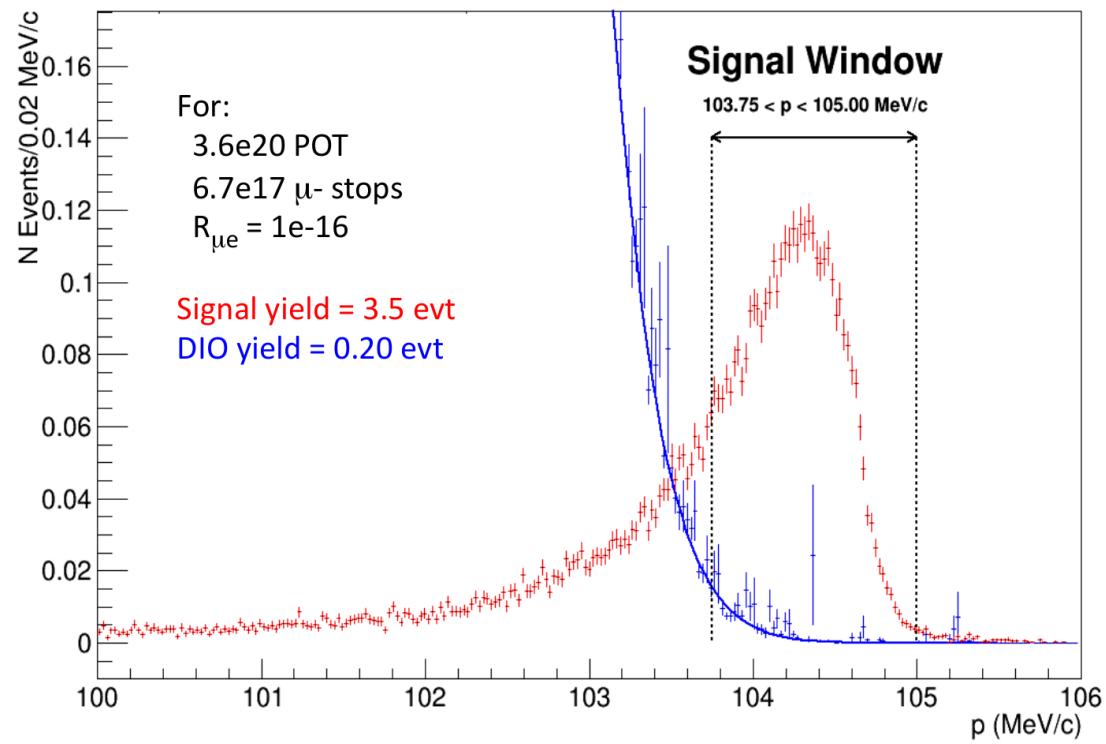
Tracking reconstruction based on  
**BABAR Kalman Filter algorithm**

No significant contribution of  
mis-reconstructed background

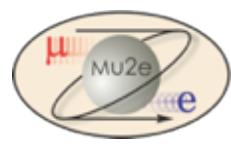
## Momentum resolution for CE

core  $\sigma \sim 120$  keV

tail  $\sigma \sim 175$  keV (2.5%)



# Mu2e Expected Background

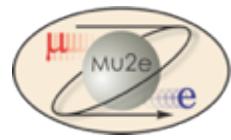


(assuming ~ 10 GHz muon stops,  $6 \times 10^{17}$  stopped muons in  $6 \times 10^7$  s of beam time)

Category	Background process	Estimated yield (events)
Intrinsic	Muon decay-in-orbit (DIO)	$0.199 \pm 0.092$
	Muon capture (RMC)	$0.000^{+0.004}_{-0.000}$
Late Arriving	Pion capture (RPC)	$0.023 \pm 0.006$
	Muon decay-in-flight ( $\mu$ -DIF)	$<0.003$
	Pion decay-in-flight ( $\pi$ -DIF)	$0.001 \pm <0.001$
Miscellaneous	Beam electrons	$0.003 \pm 0.001$
	Antiproton induced	$0.047 \pm 0.024$
	Cosmic ray induced	$0.092 \pm 0.020$
Total		$0.37 \pm 0.10$

**Discovery sensitivity accomplished by suppressing backgrounds to  $< 0.5$  event total**

**Upper Limit  $< 6 \times 10^{-17}$  @ 90% C.L.**



## Project-X re-imagined to match Budget constraints:

### 1) PIP-2 plans:

- 1 MW at LBNF at start (2025)
- 2 MW at regime at LBNF
- $\times 10$  at Mu2e

[Projectx-docdb.fnal.gov/cgi-bin/  
ShowDocument?docid=1232](http://Projectx-docdb.fnal.gov/cgi-bin/ShowDocument?docid=1232)  
CLVF-snowmass → Arxiv.1311.5278  
Mu2e-2 → Arxiv.1307.1168v2.pdf

### 2) Depending on the beam Structure available:

- study Z dependence  
if signal is observed

### 3) If no signal is observed

Use  $\times 10$  events in Mu2e-2

Minor modifications of the  
detector →  $\text{BR} < 6 \times 10^{-18}$

*V. Cirigliano, R. Kitano, Y. Okada, P. Tuzon, arXiv:0904.0957 [hep-ph],  
Phys. Rev. D80 (2009) 013002*

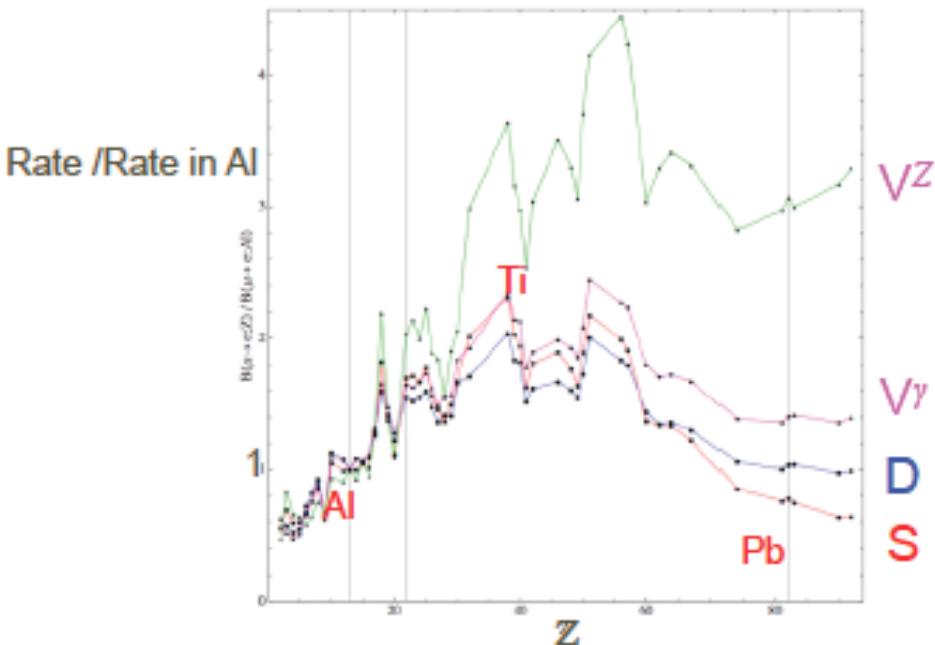


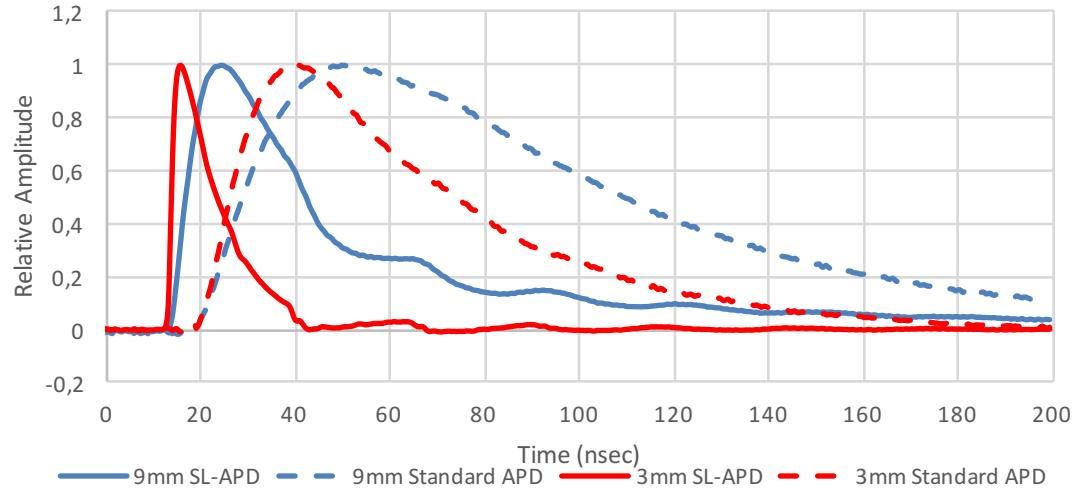
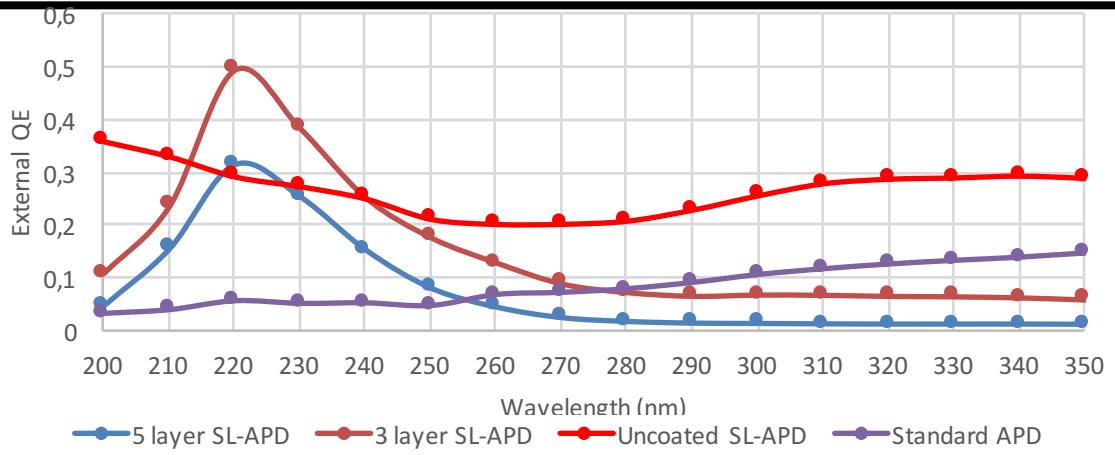
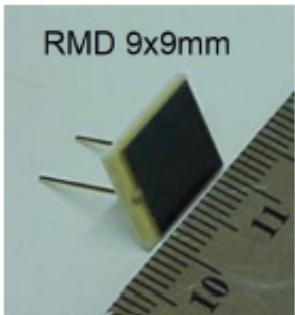
Figure 3: Target dependence of the  $\mu \rightarrow e$  conversion rate in different single-operator dominance models. We plot the conversion rates normalized to the rate in Aluminum ( $Z = 13$ ) versus the atomic number  $Z$  for the four theoretical models described in the text:  $D$  (blue),  $S$  (red),  $V^{(\gamma)}$  (magenta),  $V^{(Z)}$  (green). The vertical lines correspond to  $Z = 13$  (Al),  $Z = 22$  (Ti), and  $Z = 83$  (Pb).

# Photosensors Choice

A Caltech/JPL/RMD consortium formed to develop a Large area RMD APD **into a super-lattice APD with high Q.E. @ 220 nm** incorporating also **an Atomic Layer Deposition antireflection filter** to reduce efficiency for wavelength > 300 nm.

- ✓ 60% QE @ 220 nm
- ✓ ~ 0.1 % QE @ 300 nm
- ✓ capacitance ~ 60 pF  
(1/5 of Ham S8664)
- ✓ HV ~ 1800 V
- ✓ Operation Gain ~ 500
- ✓ Decay time ~ 25 ns.

deltadoped APD from RMD



# CsI Crystals

	LYSO	BaF <sub>2</sub>	CsI
Radiation Length X <sub>0</sub> [cm]	1.14	2.03	1.86
Light Yield [% NaI(Tl)]	75	4/36	3.6
Decay Time[ns]	40	0.9/650	20
Photosensor	APD	R&D APD	SiPM
Wavelength [nm]	402	220/300	310

## LYSO

CDR

- Radiation hard, not hygroscopic
- Excellent LY
- Tau = 40ns
- Emits @ 420 nm,
- Easy to match to APD.
- High cost > 40\$/cc

## Barium Fluoride (BaF<sub>2</sub>)

BASELINE-TDR

- Radiation hard, not hygroscopic
- very fast (220 nm) scintillating light
- Larger slow component at 300 nm. should be suppress for high rate capability
- Photo-sensor should have extended UV sensitivity and be "solar"-blind
- Medium cost 10\$/cc

## CsI(pure)

Baseline for EDR

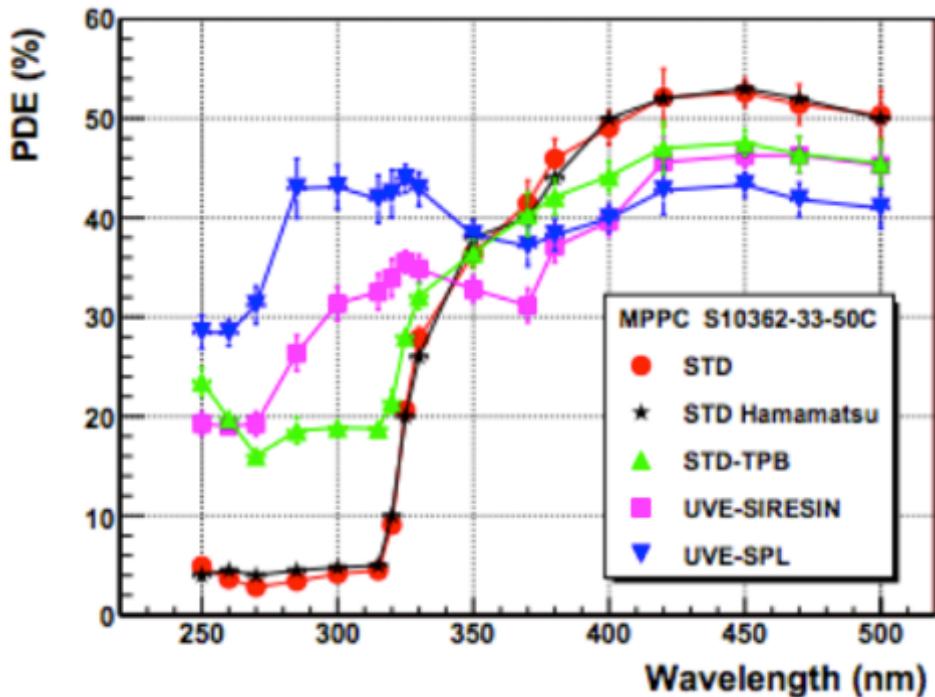
- Not too radiation hard
- Slightly hygroscopic
- 15-20 ns emission time
- Emits @ 320 nm.
- Comparable LY of fast component of BaF<sub>2</sub>.
- Cheap (6-8 \$/cc)

# UV extended SiPM

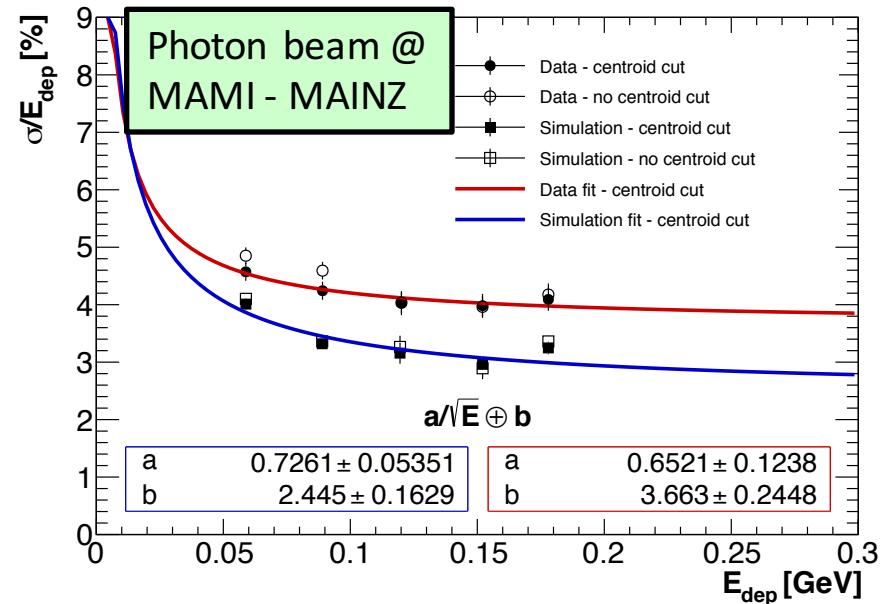
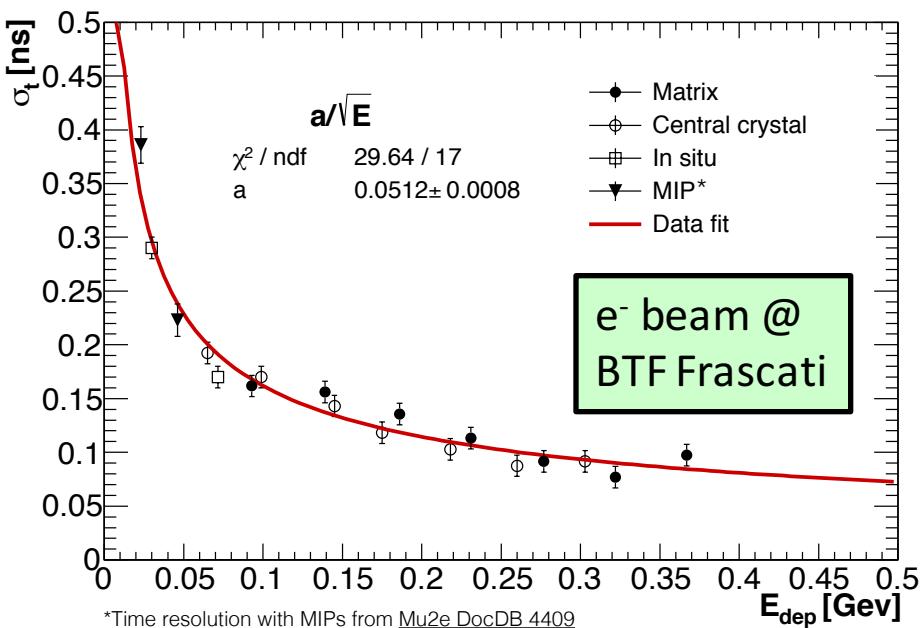
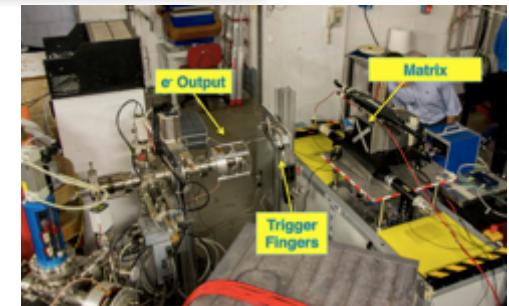
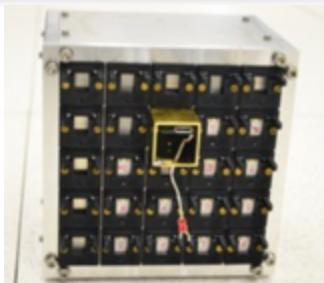
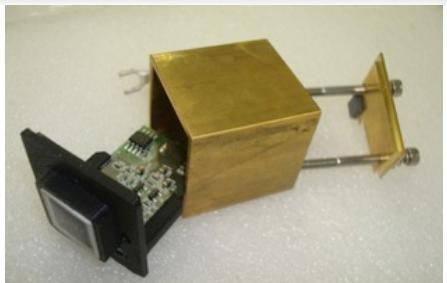
The basic SiPM element (pixel) is a combination of the Geiger-APD and quenching resistor

**The PDE of UV-enhanced MPPC is higher than the standard one:**

- **30-40% @ 310 nm** (CsI pure wavelength)
- with new silicon resin window
- Gain  $\sim 10^6$



# LYSO Legacy



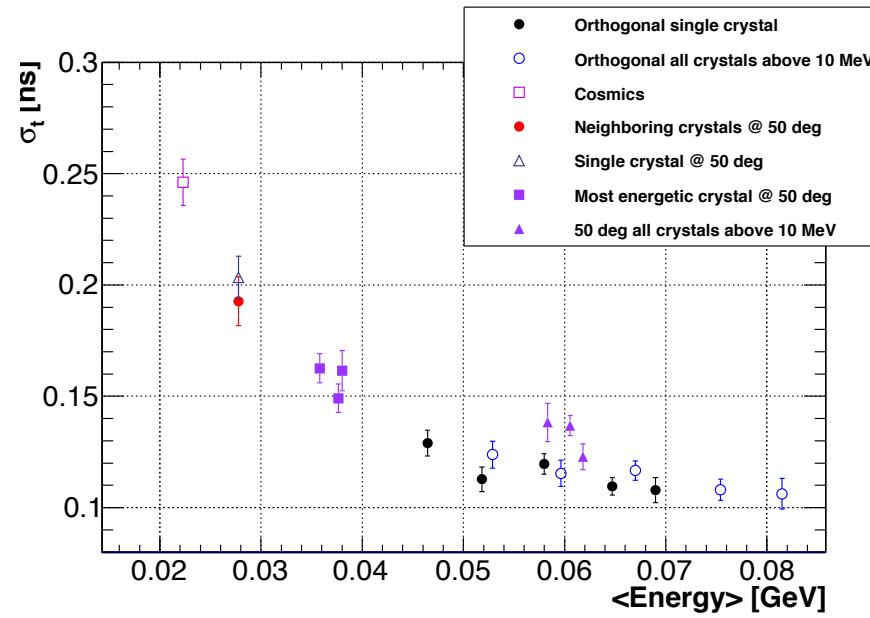
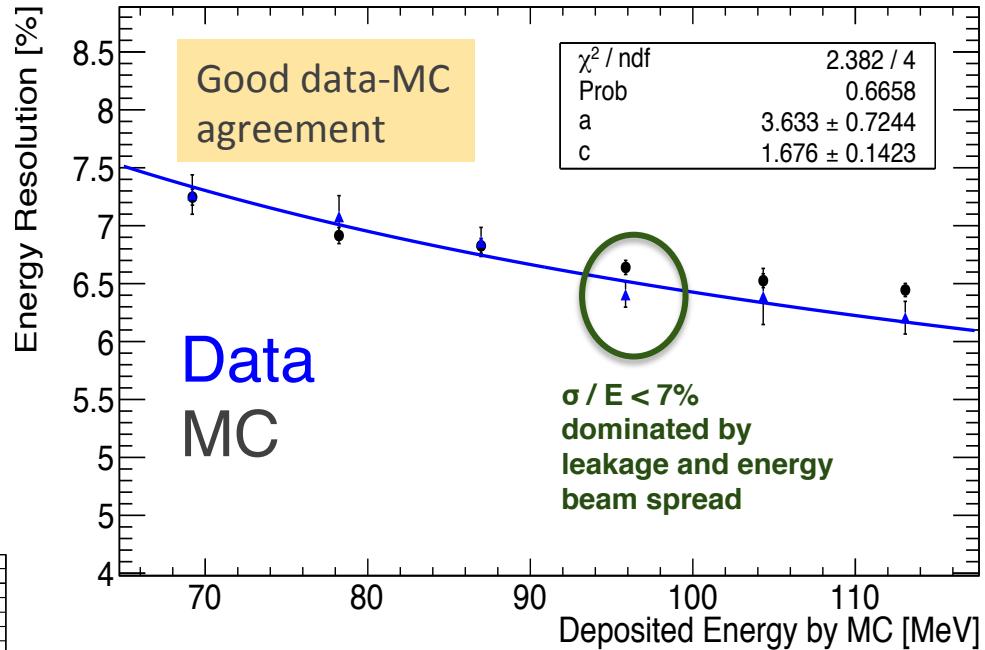
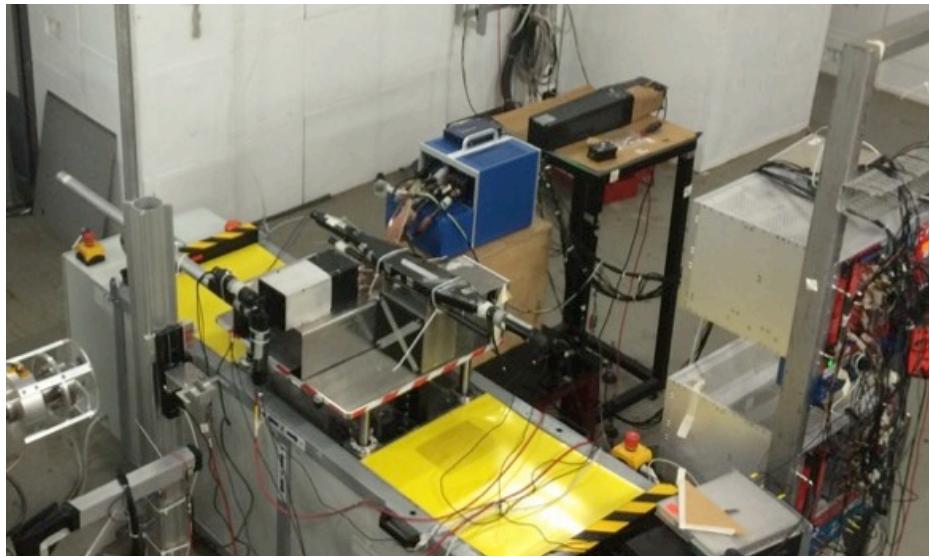
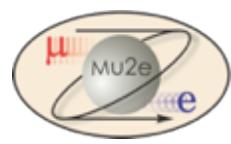
$\sigma_T = 51 \text{ ps}/\sqrt{E/\text{GeV}}$   
compare with KLOE  
 $\sim 55 \text{ ps}/\sqrt{E/\text{GeV}}$

Energy resolution as a function of the energy deposition fitted with the function:

$\sim 4\% @ 100 \text{ MeV}$   $\frac{\sigma_E}{E} = \frac{a}{\sqrt{E}} \oplus \frac{b}{E} \oplus c$

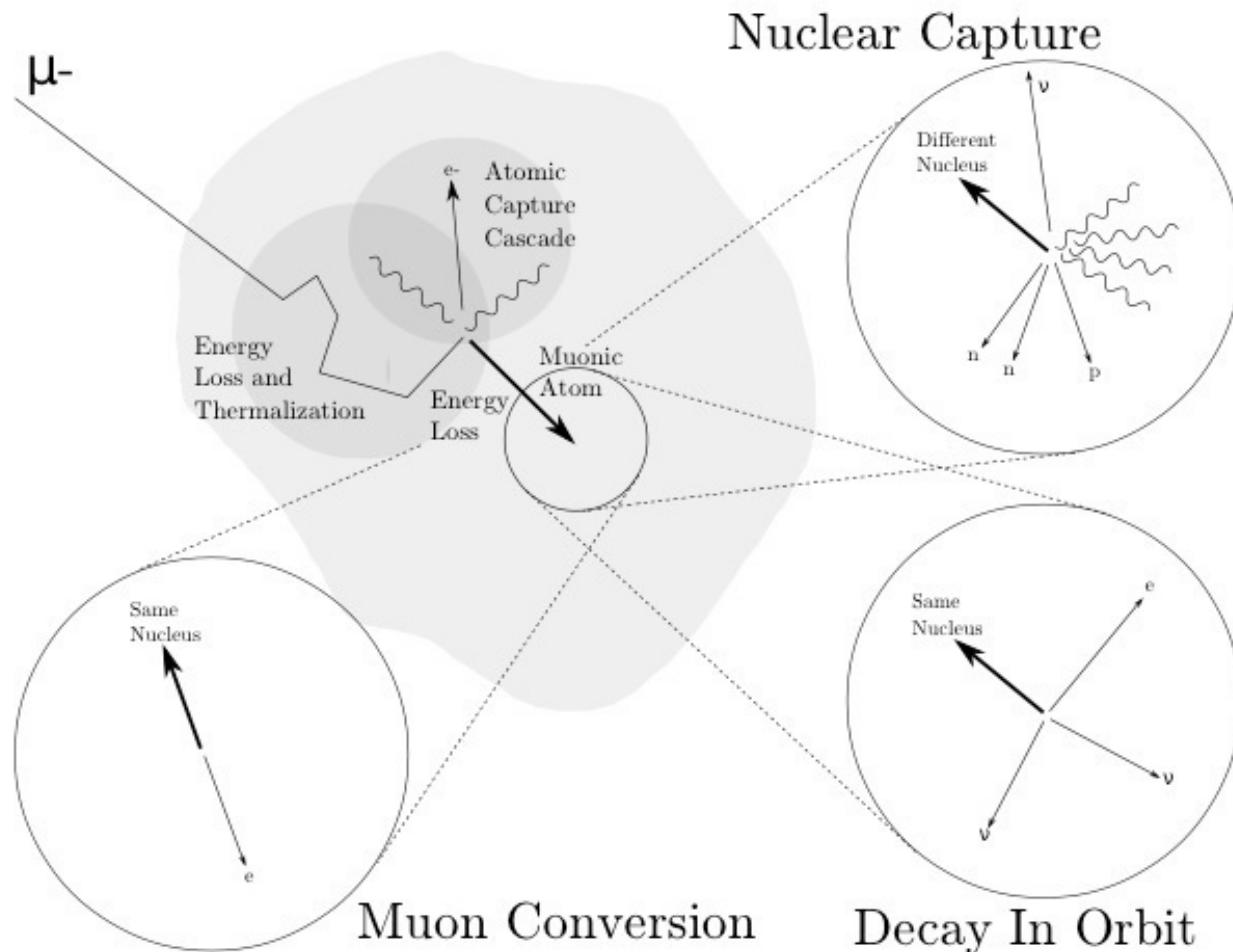
Noise term  $b$  considered negligible ( $\sim 0.1\%$  in quadrature).

# CsI+MPPC backup option

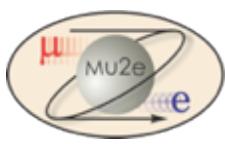


# Muon Processes

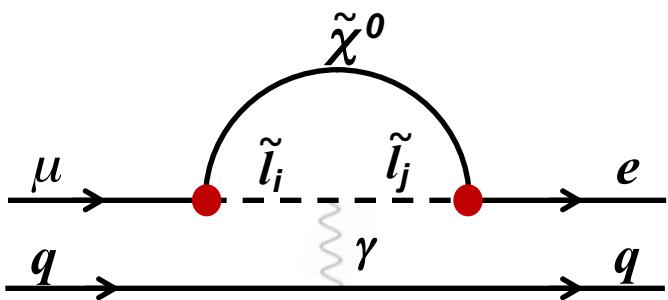
The atomic, nuclear, and particle physics of  $\mu^-$  drive the design of the experiment



# Specific Example: SUSY



## Probe SUSY through loops

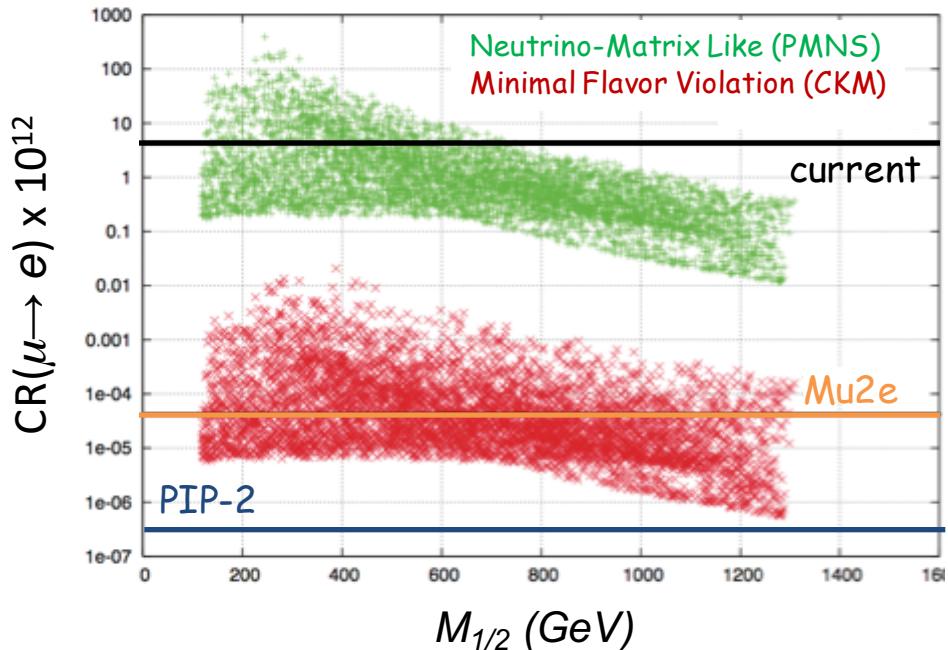


If SUSY seen at LHC  $\rightarrow$  rate  $\sim 10^{-15}$

Implies  $\sim 40\text{-}50$  signal events with negligible background in Mu2e for many SUSY models.

## SUSY GUT in an SO(10) framework

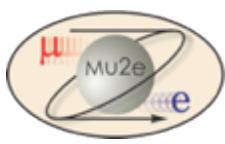
$\mu N \rightarrow e N$  ( $\tan\beta = 10$ )



L. Calibbi et al., [hep-ph/0605139](#)

**Complementary with the LHC experiments  
while providing models' discrimination**

# Other CLFV Predictions



M.Blanke, A.J.Buras, B.Duling, S.Recksiegel, C.Tarantino

ratio	LHT	MSSM (dipole)	MSSM (Higgs)
$\frac{Br(\mu^- \rightarrow e^- e^+ e^-)}{Br(\mu \rightarrow e\gamma)}$	0.02...1	$\sim 6 \cdot 10^{-3}$	$\sim 6 \cdot 10^{-3}$
$\frac{Br(\tau^- \rightarrow e^- e^+ e^-)}{Br(\tau \rightarrow e\gamma)}$	0.04...0.4	$\sim 1 \cdot 10^{-2}$	$\sim 1 \cdot 10^{-2}$
$\frac{Br(\tau^- \rightarrow \mu^- \mu^+ \mu^-)}{Br(\tau \rightarrow \mu\gamma)}$	0.04...0.4	$\sim 2 \cdot 10^{-3}$	0.06...0.1
$\frac{Br(\tau^- \rightarrow e^- \mu^+ \mu^-)}{Br(\tau \rightarrow e\gamma)}$	0.04...0.3	$\sim 2 \cdot 10^{-3}$	0.02...0.04
$\frac{Br(\tau^- \rightarrow \mu^- e^+ e^-)}{Br(\tau \rightarrow \mu\gamma)}$	0.04...0.3	$\sim 1 \cdot 10^{-2}$	$\sim 1 \cdot 10^{-2}$
$\frac{Br(\tau^- \rightarrow e^- e^+ e^-)}{Br(\tau^- \rightarrow e^- \mu^+ \mu^-)}$	0.8...2.0	$\sim 5$	0.3...0.5
$\frac{Br(\tau^- \rightarrow \mu^- \mu^+ \mu^-)}{Br(\tau^- \rightarrow \mu^- e^+ e^-)}$	0.7...1.6	$\sim 0.2$	5...10
$\frac{R(\mu Ti \rightarrow e Ti)}{Br(\mu \rightarrow e\gamma)}$	$10^{-3} \dots 10^2$	$\sim 5 \cdot 10^{-3}$	0.08...0.15

arXiv:0909.5454v2[hep-ph]

Table 3: Comparison of various ratios of branching ratios in the LHT model ( $f = 1$  TeV) and in the MSSM without [92, 93] and with [96, 97] significant Higgs contributions.

- Relative rates are model dependent
- Measure ratios to pin-down theory details

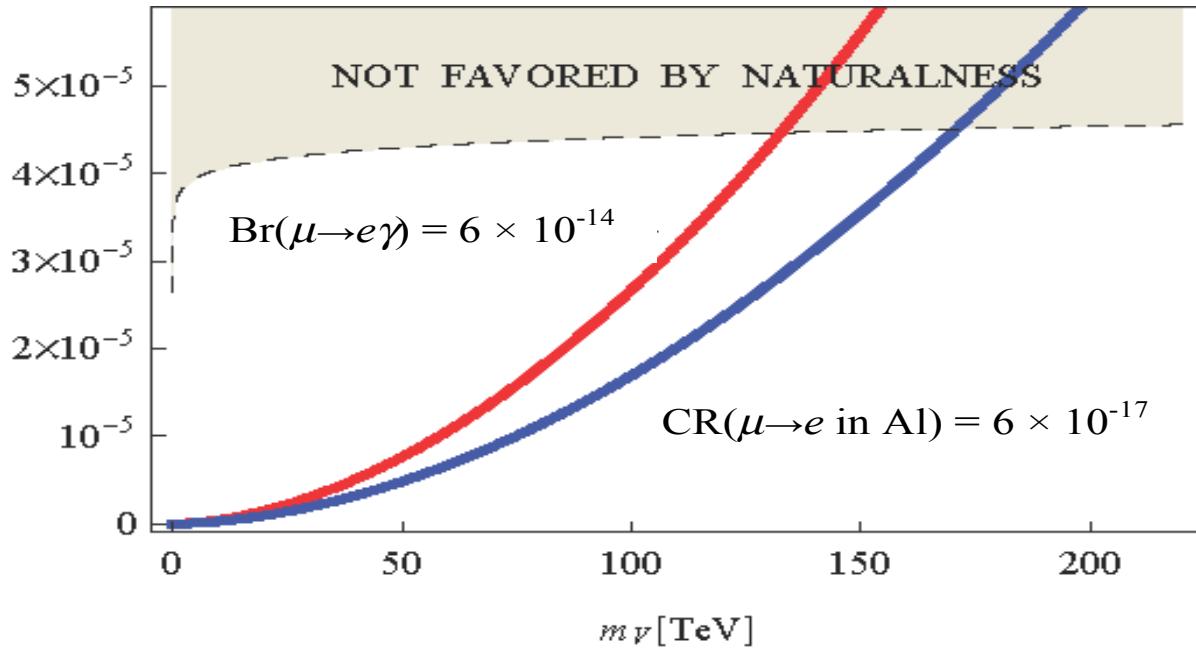
# SUSY benchmark points vs LHC

TABLE XII: LFV rates for points **SPS 1a** and **SPS 1b** in the CKM case and in the  $U_{e3} = 0$  PMNS case. The processes that are within reach of the future experiments (MEG, SuperKEKB) have been highlighted in boldface. Those within reach of post-LHC era planned/discussed experiments (PRISM/PRIME, Super Flavour factory) highlighted in italics.

Process	SPS 1a		SPS 1b		SPS 2		SPS 3		Future Sensitivity
	CKM	$U_{e3} = 0$	CKM	$U_{e3} = 0$	CKM	$U_{e3} = 0$	CKM	$U_{e3} = 0$	
$\text{BR}(\mu \rightarrow e \gamma)$	<b><math>3.2 \cdot 10^{-14}</math></b>	<b><math>3.8 \cdot 10^{-13}</math></b>	<b><math>4.0 \cdot 10^{-13}</math></b>	<b><math>1.2 \cdot 10^{-12}</math></b>	$1.3 \cdot 10^{-15}$	$8.6 \cdot 10^{-15}$	$1.4 \cdot 10^{-15}$	$1.2 \cdot 10^{-14}$	$\mathcal{O}(10^{-14})$
$\text{BR}(\mu \rightarrow e e e)$	$2.3 \cdot 10^{-16}$	$2.7 \cdot 10^{-15}$	$2.9 \cdot 10^{-16}$	$8.6 \cdot 10^{-15}$	$9.4 \cdot 10^{-18}$	$6.2 \cdot 10^{-17}$	$1.0 \cdot 10^{-17}$	$8.9 \cdot 10^{-17}$	$\mathcal{O}(10^{-14})$
$\text{CR}(\mu \rightarrow e \text{ in Ti})$	$2.0 \cdot 10^{-15}$	$2.4 \cdot 10^{-14}$	$2.6 \cdot 10^{-15}$	$7.6 \cdot 10^{-14}$	$1.0 \cdot 10^{-16}$	$6.7 \cdot 10^{-16}$	$1.0 \cdot 10^{-16}$	$8.4 \cdot 10^{-16}$	$\mathcal{O}(10^{-18})$
$\text{BR}(\tau \rightarrow e \gamma)$	$2.3 \cdot 10^{-12}$	$6.0 \cdot 10^{-13}$	$3.5 \cdot 10^{-12}$	$1.7 \cdot 10^{-12}$	$1.4 \cdot 10^{-13}$	$4.8 \cdot 10^{-15}$	$1.2 \cdot 10^{-13}$	$4.1 \cdot 10^{-14}$	$\mathcal{O}(10^{-8})$
$\text{BR}(\tau \rightarrow e e e)$	$2.7 \cdot 10^{-14}$	$7.1 \cdot 10^{-15}$	$4.2 \cdot 10^{-14}$	$2.0 \cdot 10^{-14}$	$1.7 \cdot 10^{-15}$	$5.7 \cdot 10^{-17}$	$1.5 \cdot 10^{-15}$	$4.9 \cdot 10^{-16}$	$\mathcal{O}(10^{-8})$
$\text{BR}(\tau \rightarrow \mu \gamma)$	$5.0 \cdot 10^{-11}$	$1.1 \cdot 10^{-8}$	$7.3 \cdot 10^{-11}$	$1.3 \cdot 10^{-8}$	$2.9 \cdot 10^{-12}$	$7.8 \cdot 10^{-10}$	$2.7 \cdot 10^{-12}$	$6.0 \cdot 10^{-10}$	$\mathcal{O}(10^{-9})$
$\text{BR}(\tau \rightarrow \mu \mu \mu)$	$1.6 \cdot 10^{-13}$	$3.4 \cdot 10^{-11}$	$2.2 \cdot 10^{-13}$	$3.9 \cdot 10^{-11}$	$8.9 \cdot 10^{-15}$	$2.4 \cdot 10^{-12}$	$8.7 \cdot 10^{-15}$	$1.9 \cdot 10^{-12}$	$\mathcal{O}(10^{-8})$

- These are SuSy benchmark points for which LHC has discovery sensitivity
- Some of these will be observable by MEG/Belle-2
- All of these will be observable by Mu2e

# Specific example: Leptoquarks



## Leptoquarks

Presenza di leptoquarks alla scala del TeV potrebbe indurre processi CLFV con una costante di accoppiamento  $\lambda$ .

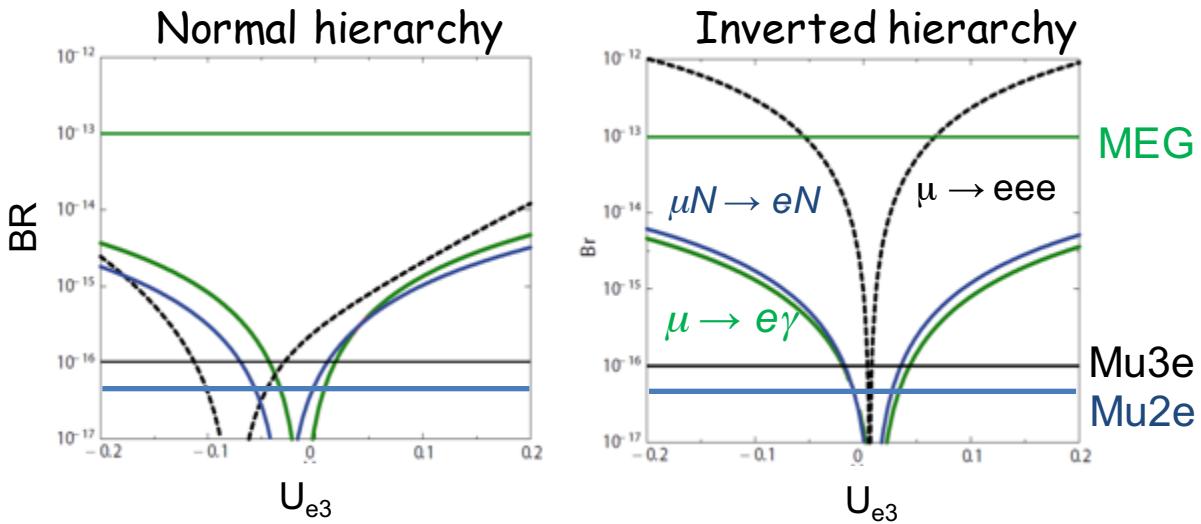
- Rosso: MEG-II
- Blu: Mu2e

# Specific example: Higgs Triplet e LHT

M. Kakizaki et al., PLB566 (2003) 210

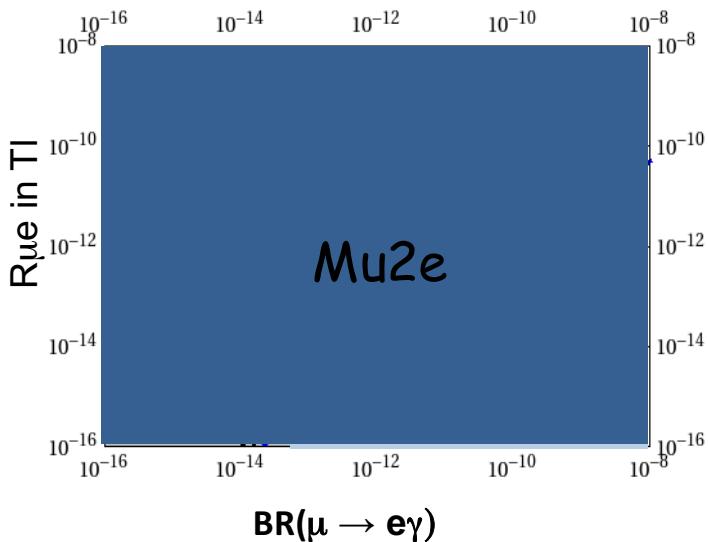
## Higgs triplet model

Dependence on  
neutrino mass  
hierarchy and  $\theta_{13}$

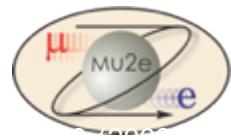


## Littlest Higgs with T-parity

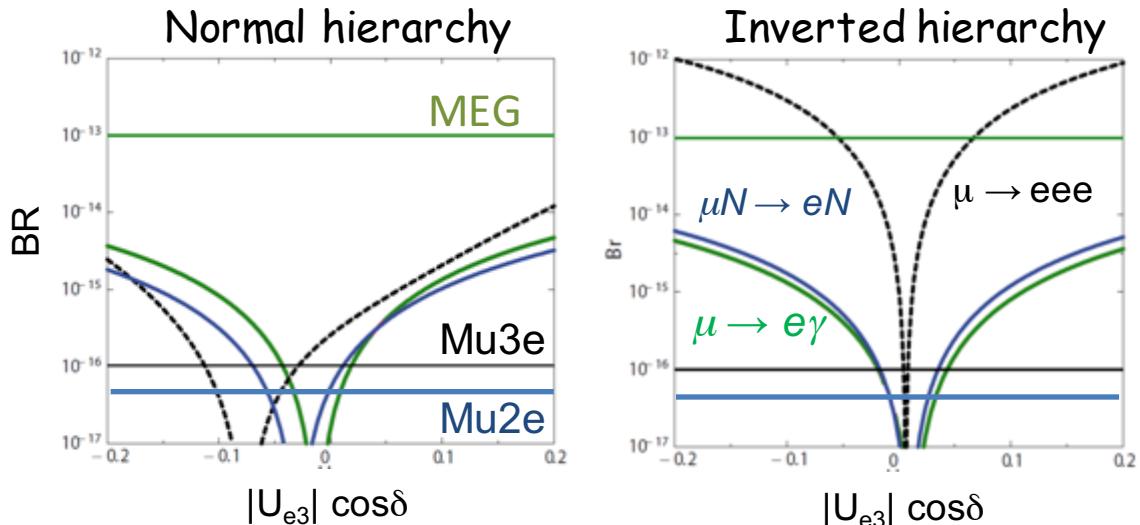
M. Blanke et al., Acta Phys.Polon.B41:657,2010



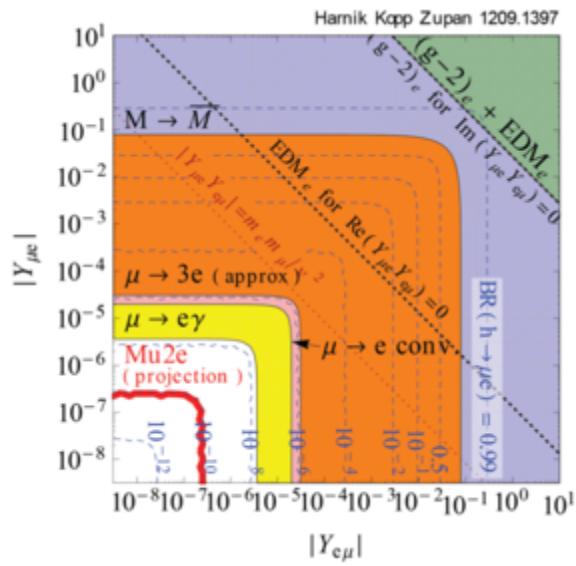
# A few more models...



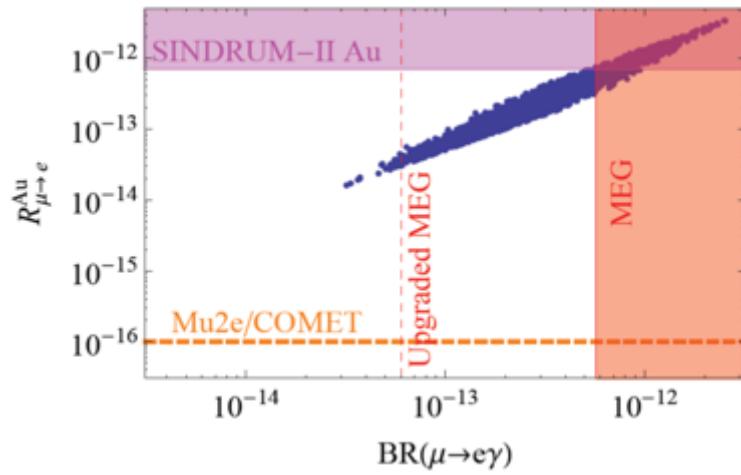
## Higgs triplet model



## Flavor violating Yukawa couplings



## Left-right symmetric models



# Stopping Target Monitor

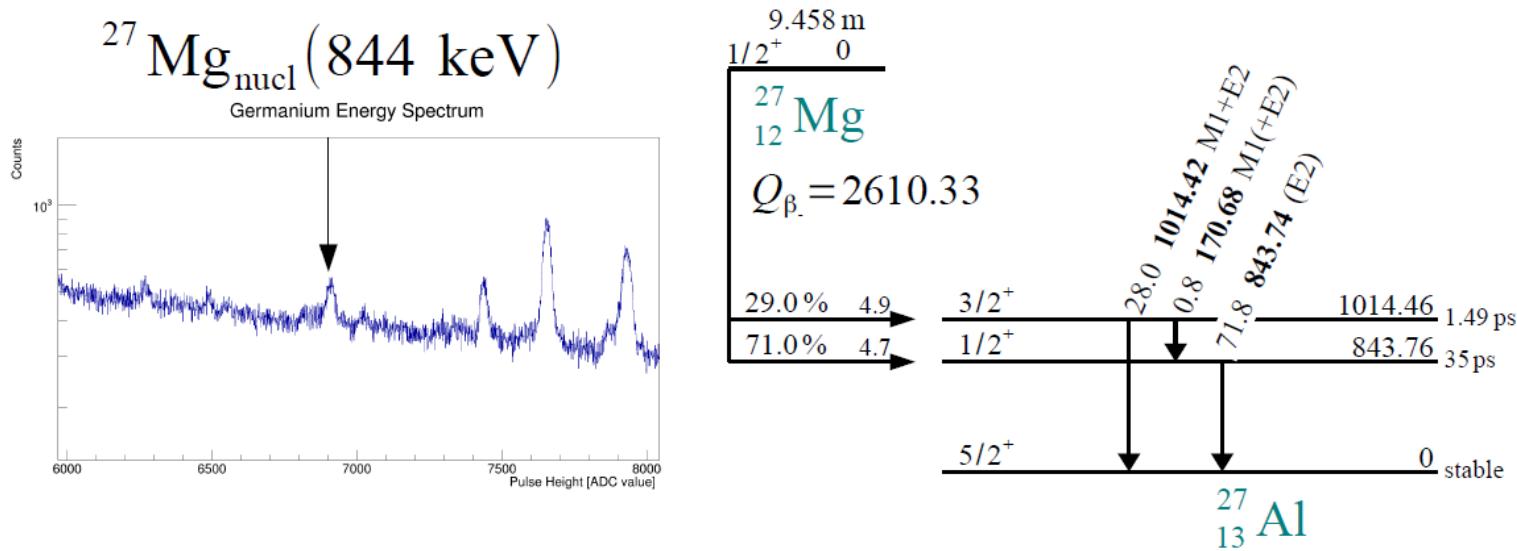


Figure 7.18. Preliminary singles germanium spectrum from the AlCap experiment at PSI. When muons stop in aluminum, they capture on the nucleus 60% of the time. A fraction of the captures produce  $^{27}\text{Mg}$  in the ground state, which has a half-life of 9.5 minutes. In the decay, an 844 keV gamma is produced 72% of the time.

- Need a high precise gamma detector (HpGe)
- Energy of gamma ray is unique to the detector
- Detecting the delayed gamma rays eliminate problems related to beam flash
- Proton beam structure is 0.5 s on followed by 0.8 s idle. Gamma spectrum will be acquired during idle time.
- Hpge should view the target far from the source and beyond DS

# CE trajectory

