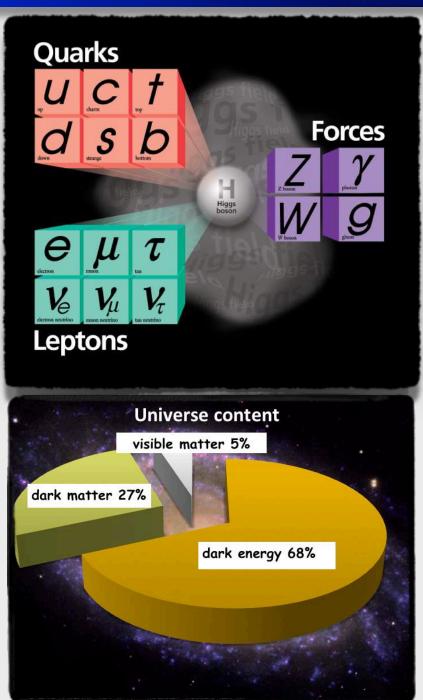


Newcomers lunch

Standard Model



- Standard Modern successful theory
- Open questions:
 - Dark energy and dark matter
 - Baryon Asymmetry
 - Hierarchy problem
 - Family structures and masses
 - **)** ...
- We can look for answers at various frontiers
 - ▶ Energy, intensity or cosmic frontiers



Intensity frontier: Muon Campus



- Intensity Frontier with muons an alternative way to look for answers
- Access New Physics at higher energies through loops
 - g-2: Measuring muon's anomalous magnetic moment to 140 ppb
 - Mu2e: Looking for ultra-rare decays



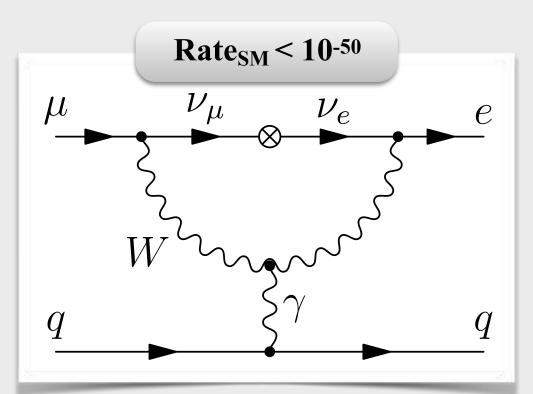
Mu2e ($\mu \rightarrow e$)

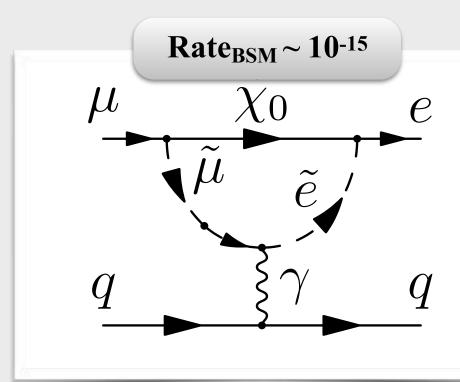


■Mu2e will search for neutrino-less, coherent muon conversion into an electron in the presence of *Al* nucleus:

$$\mu^- + Al \rightarrow e^- + Al$$

- ■In the SM, $\mu \rightarrow e^{-}$ occurs at the rate of <10⁻⁵⁰
 - Signal observation at Mu2e is unambiguous sign of new physics





Other CLFV Processes



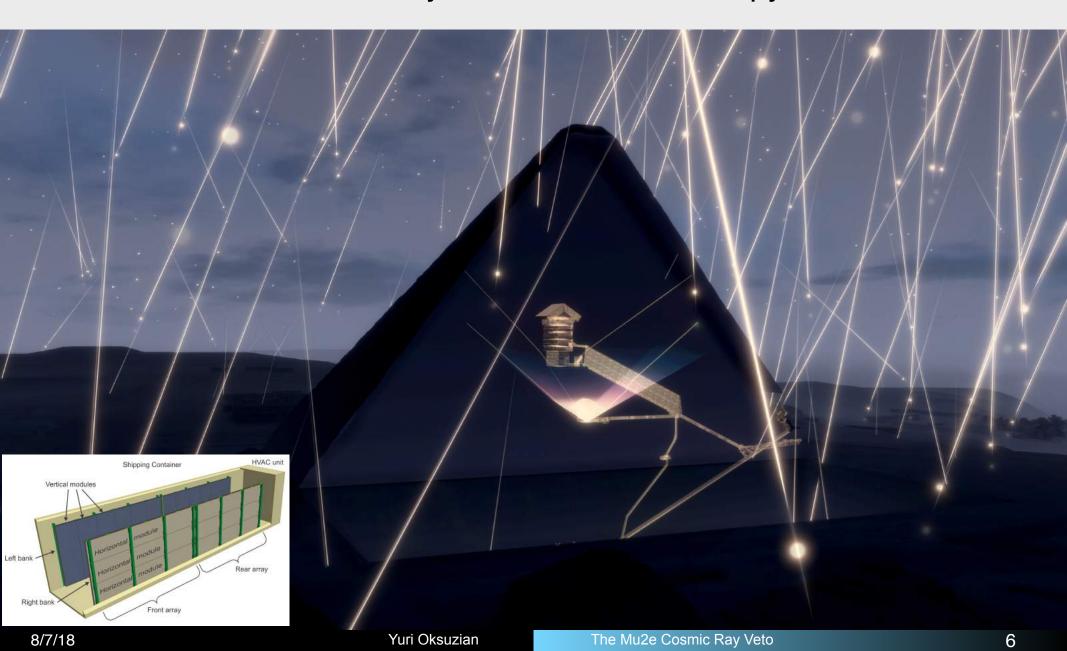
■ Most stringent limits come from muon sector

Process	Current Limit	Next Generation Exp
$ au ightarrow \mu \eta$	BR < 6.5E-8	
$ au ightarrow \mu \gamma$	BR < 6.8E-8	10 ⁻⁹ - 10 ⁻¹⁰ (Belle II)
$ au ightarrow \mu \mu \mu$	BR < 3.2E-8	
$\tau \rightarrow eee$	BR < 3.6E-8	
$K_L \rightarrow e \mu$	BR < 4.7E-12	
$K^+ \rightarrow \pi^+ e^- \mu^+$	BR < 1.3E-11	
$B^0 \rightarrow e\mu$	BR < 7.8E-8	
$B^+ \rightarrow K^+ e \mu$	BR < 9.1E-8	
$\mu^+ \rightarrow e^+ \gamma$	BR < 4.2E-13	10-14 (MEG)
$\mu^+ \rightarrow e^+ e^+ e^-$	BR < 1.0E-12	10-16 (PSI)
μ -N \rightarrow e-N	$R_{\mu e} < 7.0E-13$	10 ⁻¹⁷ (Mu2e, COMET)

Egypt pyramids tomography



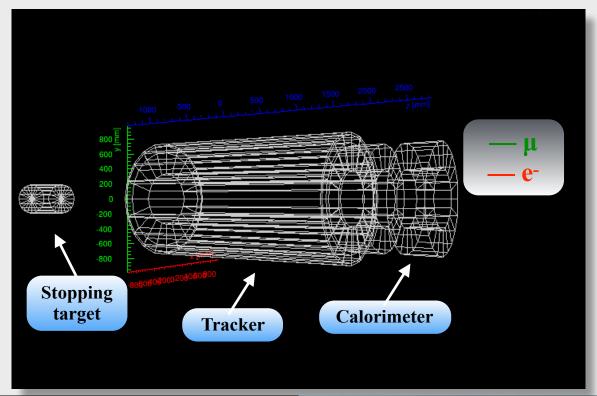
■ Use cosmic muons to study internal structures of pyramids



Cosmic Ray Background



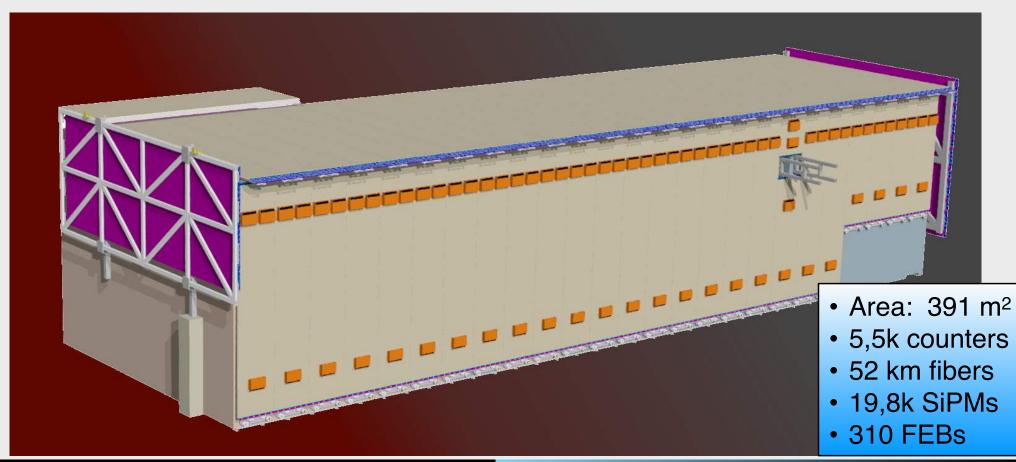
- Nasty side of muons: Mu2e is exposed to 1.25E9 cosmic muons per day
- Cosmic ray muons produce a background
 - Interactions with material inside DS, decays and faking electrons
- Mu2e expects 1 signal-like event per day induced by cosmic rays
 - 780 events over Mu2e lifetime
 - Total expected background at Mu2e: 0.41
- CRV needs to suppress cosmic ray background by a factor of 10,000



Cosmic Ray Veto



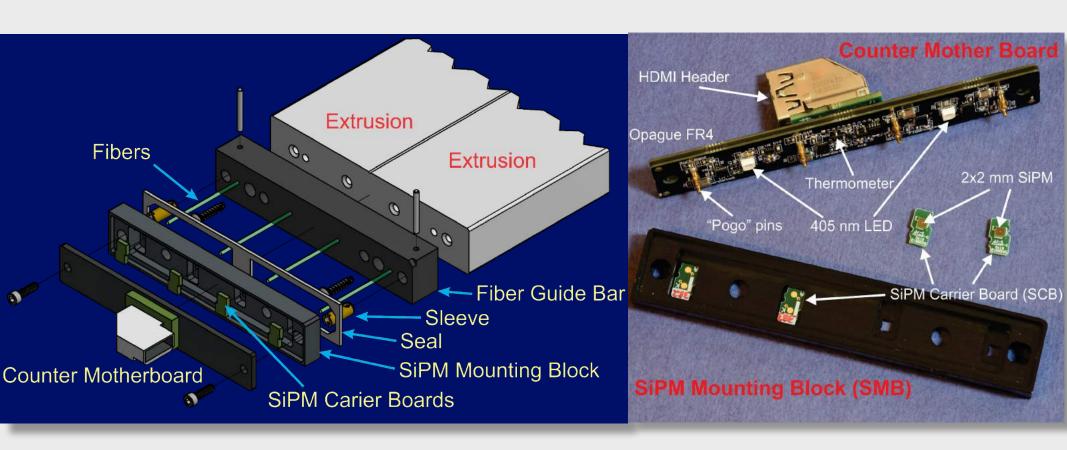
- CRV consists of 4-layer scintillating 5x2 cm² counters, read-out through wavelength-shifting fibers by 2x2 mm² SiPMs
- Cosmic ray muon detection hits coincidence in at least 3 layers localized in time and space
- Veto (offline) 125 ns from a signal window after a coincidence in the CRV

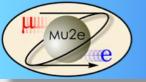


CRV building block: di-counter



- Counters: extruded PS doped with 1%PPO+0.05%POPOP, coated with TiO2
 - Glued in pairs to make di-counters
- Each counter has two 1.4 mm wavelength-shifting fibers placed in channels
- Fiber Guide Bar is glued and fly-cut
- Fibers are read out by SiPMs





Electronics



- Surface mounted SiPMs -> Carrier Boards (19,840)
- Carried Boards -> Counter Motherboards (4960)
- Counter Motherboards -> Front-end Boards (310)
- Front-end Boards -> Readout Controllers (15)

SiPM pixels: 1584 Dynamic range: 2048

Sampling rate: 12.6 ns

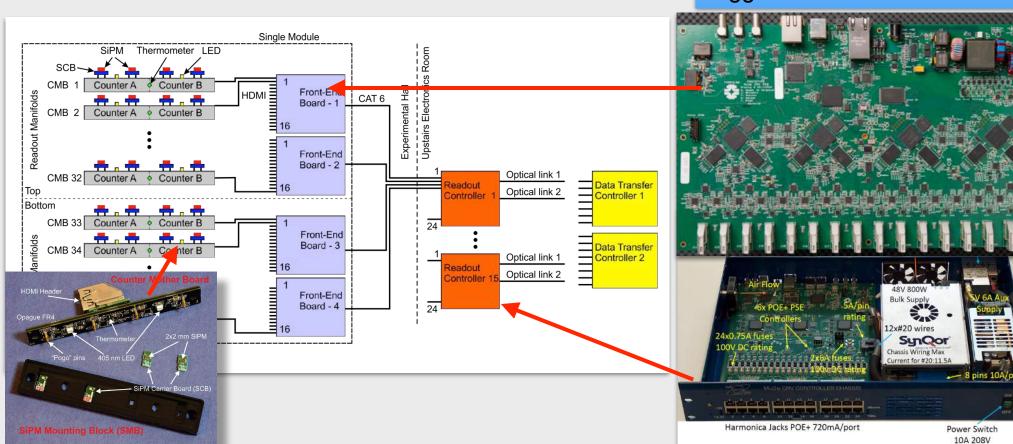
Rate to DAQ: 55 MB/s

Data per run: 0.5 PB

Time resolution: ~ 2 ns

Zero suppression: 6 PE

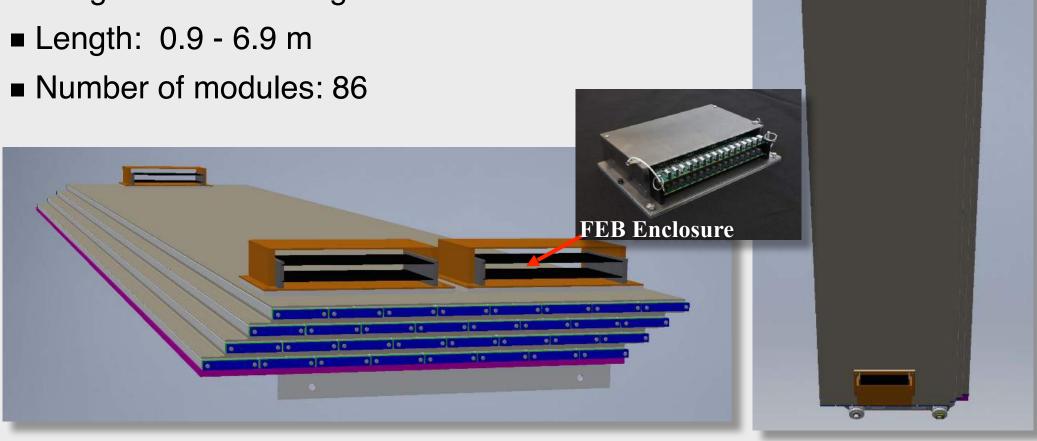
Triggered read-out



CRV module



- Di-counters are used to build a module
 - ▶ 4 layers of 8 di-counters separated by aluminum absorbers
- Entire assembly glued together
- Weight: 149 1091 kg



Module fabrication at UVa



- Module factory has been set up at University of Virginia
- 9 prototype modules have been fabricated
- Manufacturing tolerances very demanding and critical
- Production is underway
 - ▶ The factory will be producing a module/week

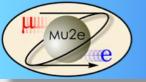


Module fabrication





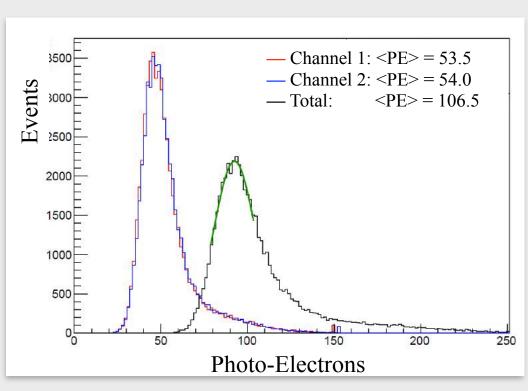
https://youtu.be/ACJTbAOXOuQ

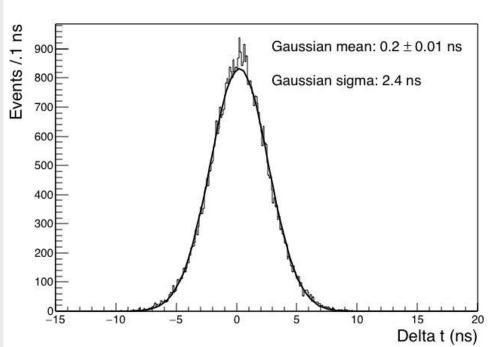


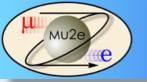
CRV performance



- Test beam results with CRV prototype
- Light yield: 54 PE/SiPM at 1 m
 - Requirement: 25 PE/SiPM
- Timing resolution: 1.7 ns
 - Requirement: 4 ns



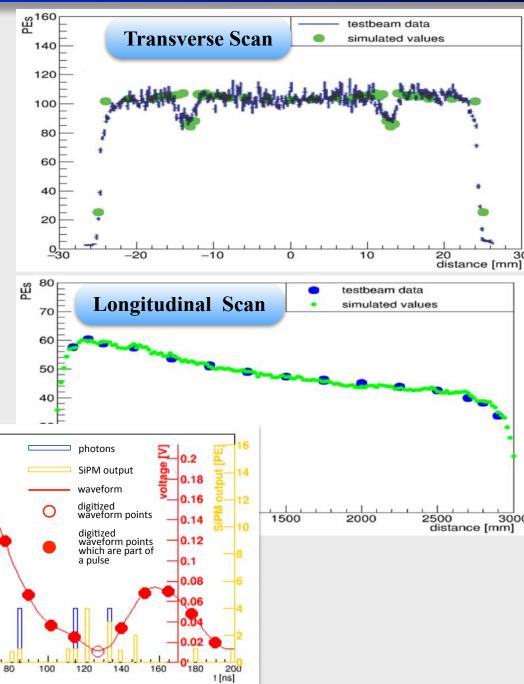


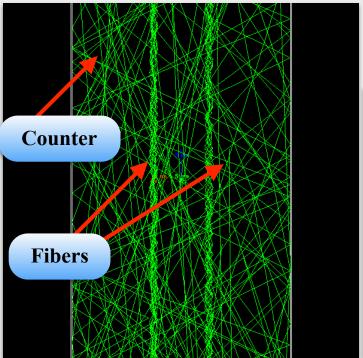


Modeling counter response



- MC has been developed to simulate CRV response to incident particles
- MC is tuned to agree with testbeam data



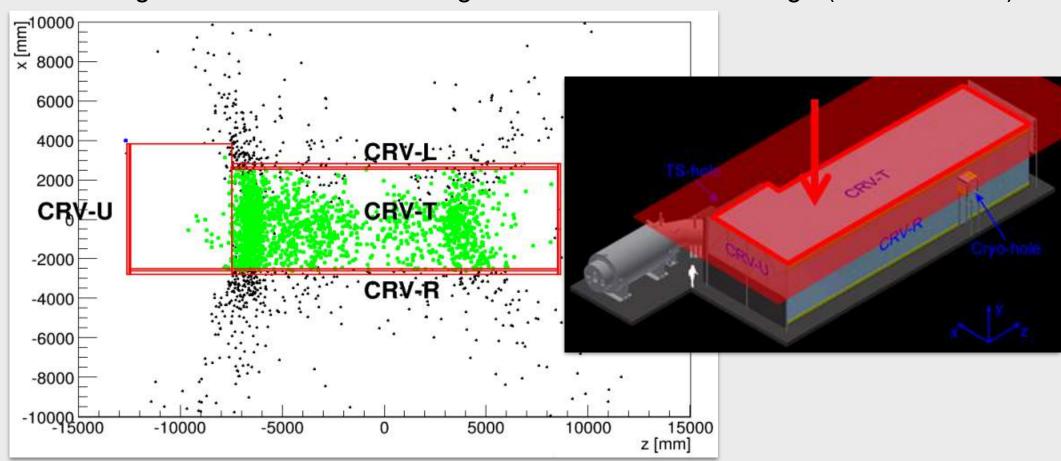


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Cosmic ray simulation



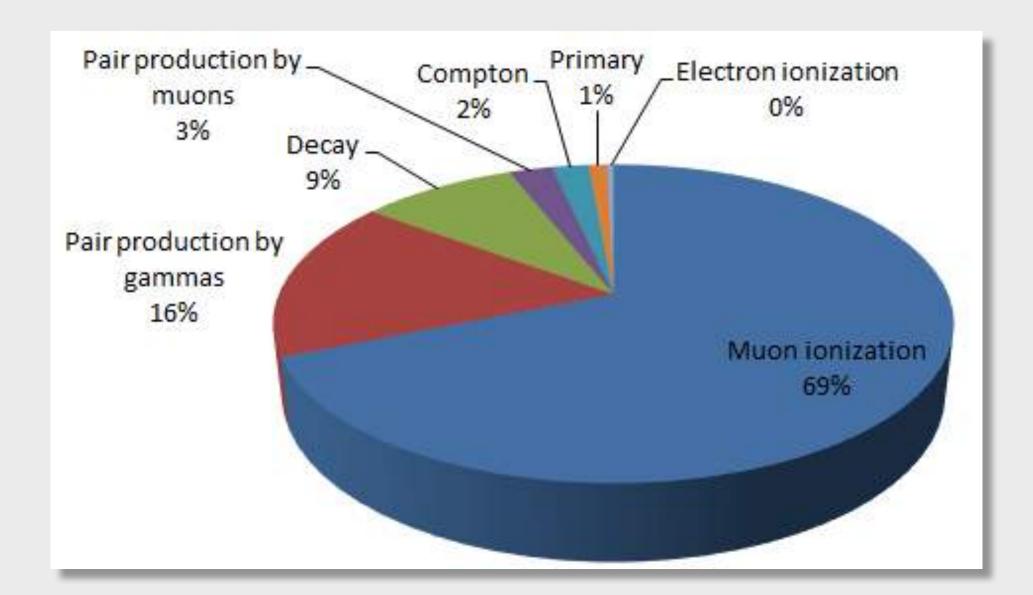
- Probability that a cosmic-ray muon produces a conversion-like event is: < 1x10⁻⁹
- Simulation campaign: 3.7x10¹² (12.8M CPU hours) cosmic ray muons generated
- Two types of simulations:
 - Global simulation: covering an area around CRV (4x live-time)
 - ▶ Targeted simulations: covering areas with limited coverage (250x live-time)



Background: Production mechanism

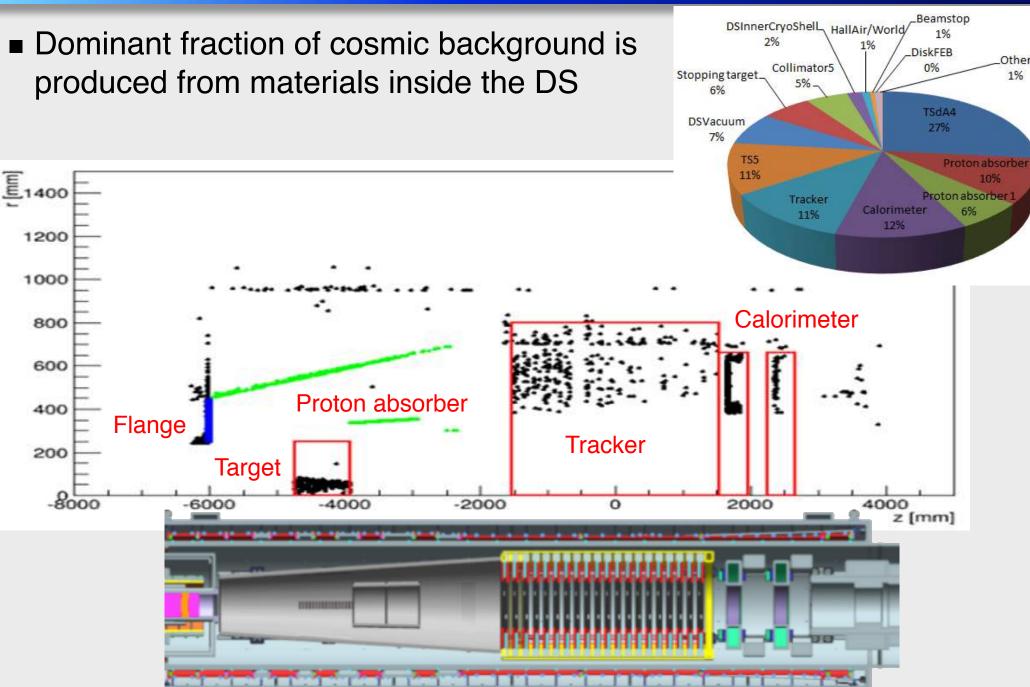


Largest source of cosmic ray background is delta-ray production



Background: Production sources

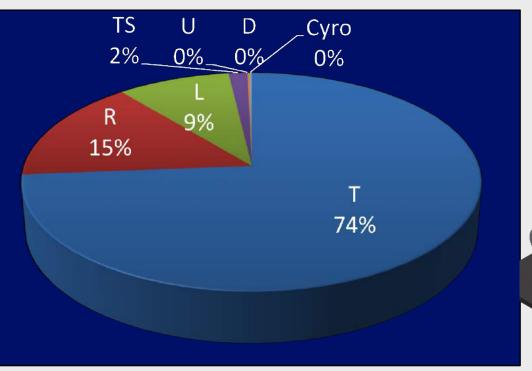


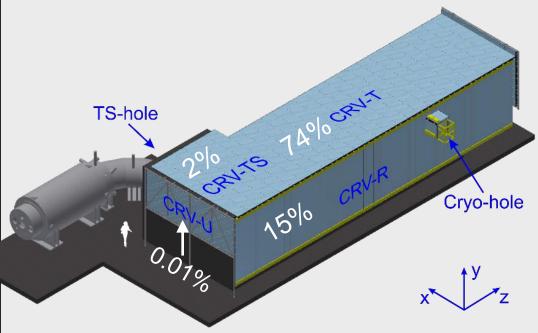


Background: Sectoral efficiencies



- CRV needs to reject 780 events over its lifetime
- Dominant (74%) fraction of muons enter from from the top
 - ▶ Top sector of CRV efficiency is >99.99%
- Other CRV regions reject much smaller fraction of muons, and require much smaller efficiencies
- Total reducible background: 0.07 events





Background: Irreducible background



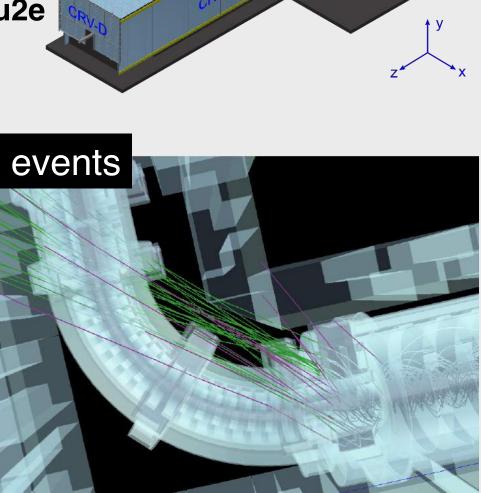
CRV-U

TS-hole

A special simulation of events going through the TS hole with 257X the total live time (770 years) we found 39 events that mimic conversion electrons

■ This is the largest background for Mu2e

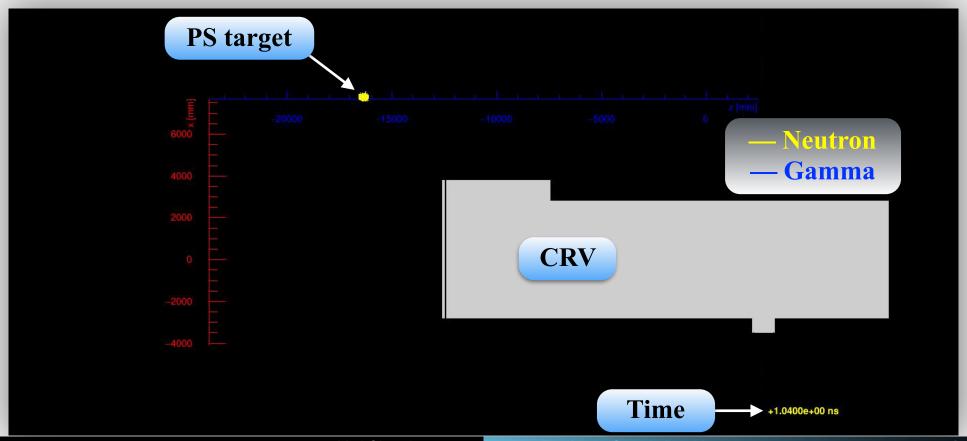
■ Two-thirds of this component can be reduced by absorbers and CRV-TS extension



Beam induced radiation

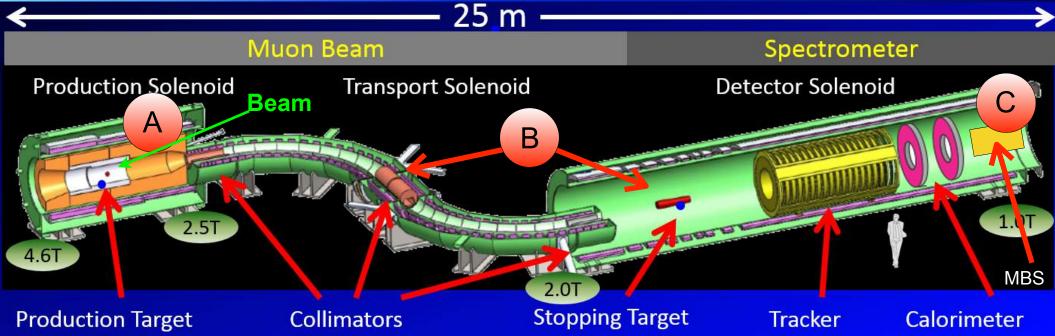


- CRV needs to operate in a sea of neutrons and gammas
- Particle fluxes from beam interactions:
 - Damage CRV components
 - Produce noise in CRV, increasing DAQ rates. Noise hits in CRV fake CR muons and increase the dead-time
 - CRV ignores hits outside of the signal window



Beam induced noise: Sources



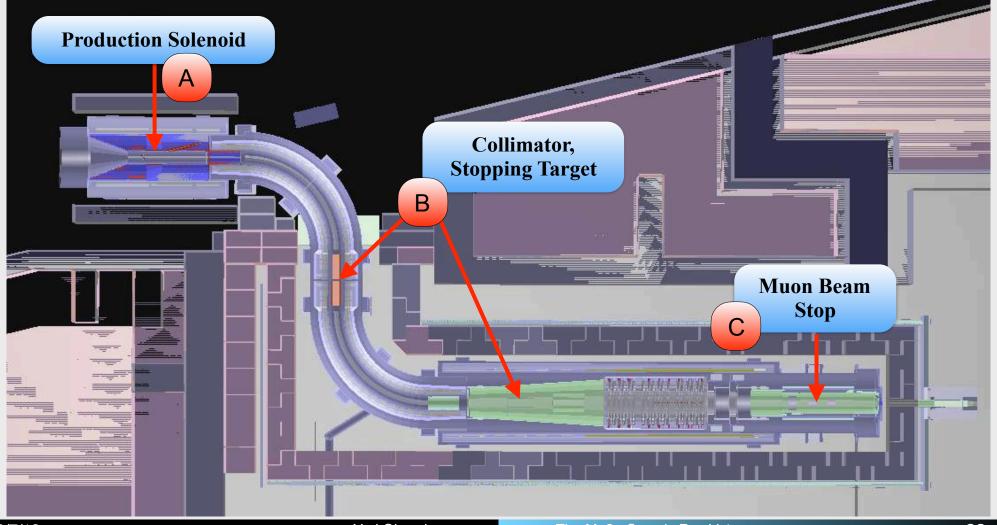


- PS is the largest source of neutrons. The source is prompt and reduced in the signal window
 - Neutrons get thermalized, captured and produce gammas
- B Fast neutrons, produced in the signal window, are from μ-captures on beam-line and stopping target
 - Fast neutron recoil off a proton depositing energy in CRV
- High energy gammas in the Muon Beam Stop (MBS): electron brems from µ-decays. Muons escaping MBS decay producing high energy electrons

CRV shielding



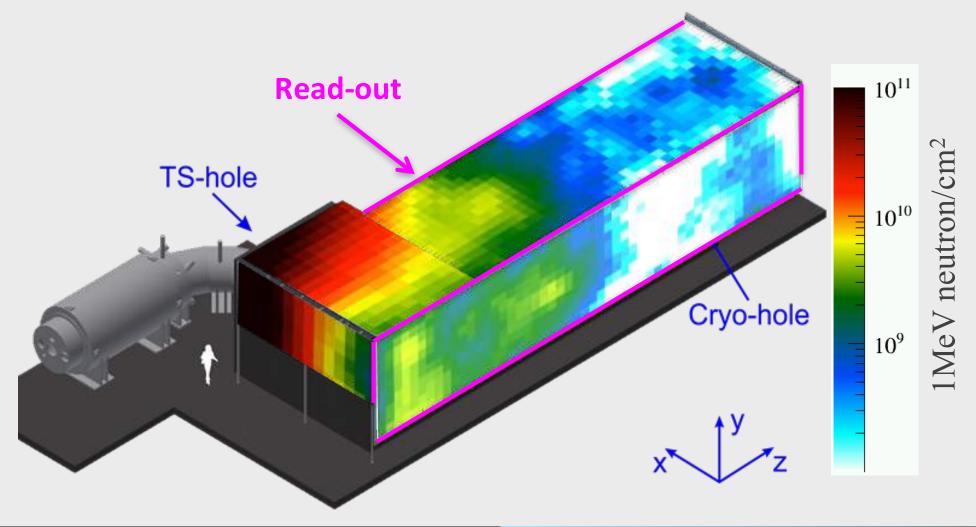
- CRV is shielded from beam induced backgrounds by 1 yd of T-shaped concrete walls
- Shielding needs to be effective and reasonably priced
 - ▶ Big effort to optimize shielding price-performance



Neutron damage to SiPM



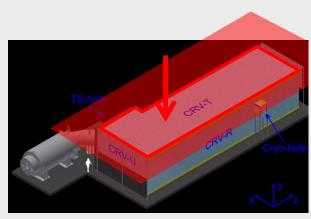
- Fast neutrons produce damage to SiPMs
 - ▶ Rad damage is driven by PS and collimators. Requirement is 10¹⁰ n/cm²
- Rad damage to scintillator and fibers is negligible

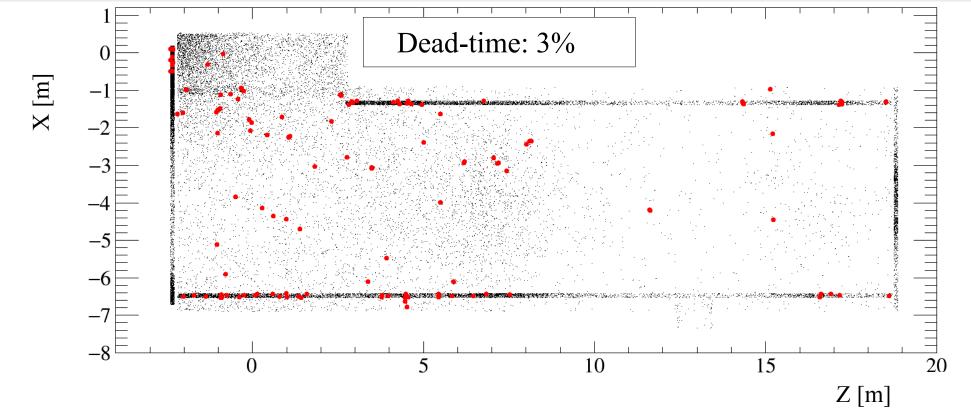


Dead-time



- Background hits in CRV fake CR muons and produce dead-time
- CRV dead-time is 3%
- Dead-time is stable under beam intensity variations and the choices of CRV thresholds





CRV Status



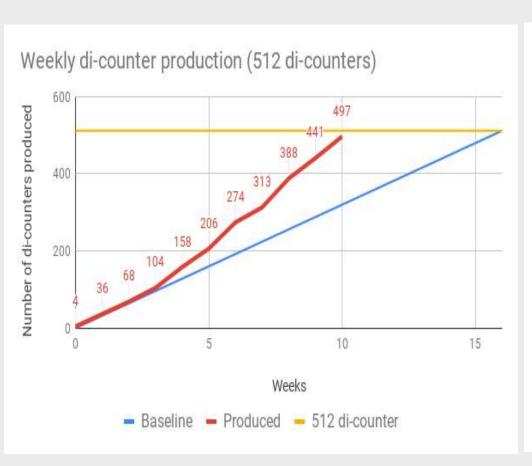


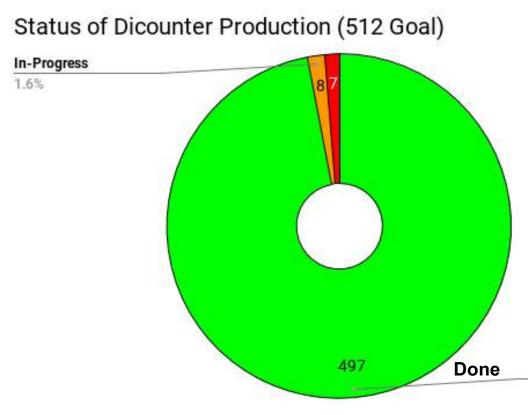


Di-counters production progress



- The goal for the Summer is to produce 512 di-counters
- We're well ahead of the production schedule
- Great effort by the team of undergraduate students





Summary



- Muons are fascinating
 - They're an excellent probe to physics beyond the SM
 - Cosmic ray muons produce a large background for Mu2e
- CRV is an essential component of Mu2e. CRV design is challenging
 - Maintain 99.99% cosmic ray veto efficiency over 3 years
 - Operate in high radiation environment
- We've just started CRV fabrication
 - The production progress is ahead of schedule
 - It will take 2 years to build CRV
- Once CRV is fabricated, it will be installed at Mu2e to shield us against cosmic muons



Backup